

# Evaluating and improving the treatment of gases in radiation schemes: the Correlated K-Distribution Model Intercomparison Project (CKDMIP)

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## 1 Introduction

The correlated k-distribution (CKD) method (e.g., Lacis and Oinas, 1991) underpins the treatment of gas absorption in the radiation schemes of most current weather and climate models. It accelerates the calculation of broadband irradiance profiles by approximating the integration over hundreds of thousands of spectral lines by a much smaller number (tens to hundreds) of quasi-monochromatic irradiance calculations, which we hereafter refer to as ‘g-points’. More g-points means higher accuracy but a larger computational cost, so we have a trade-off to make depending on the application. For climate modelling we require schemes that can accurately compute the radiative forcing of a number of different greenhouse gases over a wide range of concentrations. For short-range weather forecasting with present-day greenhouse gas concentrations, the priority is much more on efficiency: the radiation scheme must be called frequently to capture the local radiative impact of evolving cloud fields, and forecasts must be delivered to customers in a timely fashion. The lower model top in many limited-area weather models also means that, in principle, fewer g-points are required to compute the heating-rate profile.

Unfortunately, the tools and know-how to make this accuracy–efficiency trade-off are available to only a handful of specialists worldwide, with the result that most atmospheric models are available with only one gas-optics configuration, which is often not optimized for the application at hand. Indeed, Hogan et al. (2017) surveyed seven models used for the same application of global weather forecasting, and reported that the total number of g-points (short-wave plus longwave) ranged from 68 to 252. This raises the

question as to whether some schemes can achieve the same accuracy with fewer g-points.

The purpose of the Correlated K-Distribution Model Intercomparison Project (CKDMIP) is to address these issues, and specifically:

1. To use benchmark line-by-line calculations to evaluate the accuracy of existing CKD models for applications spanning short-range weather forecasting to climate modelling, and to explore how accuracy varies with number of g-points in individual CKD schemes.
2. To understand how different choices in way that CKD models are generated affects their accuracy for the same number of g-points.
3. To provide freely available datasets and software to facilitate the development of new gas-optics models, with the ultimate aim of producing a community tool to allow users to generate their own gas-optics models targeted at specific applications.

The project has similarities to the Radiative Forcing Model Intercomparison Project (RFMIP; Pincus et al., 2016), which used line-by-line calculations to evaluate the radiation schemes of a number of climate models in terms of surface and top-of-atmosphere (TOA) irradiances for a range of atmospheric profiles and climate scenarios. However, CKDMIP goes further in that it includes weather forecasting applications, and provides the means to improve the way that CKD schemes make the trade-off between accuracy and efficiency. This is possible by making available the spectral optical depth of each layer of the atmosphere due to each gas separately. The CKDMIP software package allows participants to combine the optical depths of the gases they are interested

in and perform radiative transfer calculations on the result, producing their own reference profiles of spectral or broadband irradiances and heating rates. Since the per-molecule absorption is (to a very good approximation) independent of concentration for all gases except water vapour, it is simple to produce reference calculations for scaled concentrations of individual gases.

This protocol paper describes the design and generation of these datasets and software, and what comparisons will be performed. Section 2 describes the overarching design decisions of CKDMIP, including which gases to include, which weather and climate applications to target, and for climate modelling which range of gas concentrations to consider. Section 3 describes in detail how the datasets are produced and how the spectral resolution has been chosen. Section 4 then describes what radiative transfer calculations are performed and the metrics that will be used to quantify errors in irradiances and heating rates.

Finally a note on terminology. Throughout this paper we define a *CKD scheme* as a software component (usually embedded within the radiation scheme of an atmospheric model) that takes as input profiles of atmospheric temperature, pressure and the concentrations of a number of gases, and outputs profiles of optical depth for each of a number of g-points. It also includes a means to compute the Planck function to use for each g-point. A *CKD model* is one configuration of a CKD scheme with a particular number of g-points, which might consist of a set of look-up tables that can be used by the CKD scheme. A *CKD tool* is a method (which may be fully automated or involve some hand-tuning) for generating individual CKD models, with some means to control the trade off between accuracy and the number of g-points.

## 2 Design of evaluation scenarios

### 2.1 Which gases?

In the shortwave, the choice of which gases to include is relatively uncontroversial: H<sub>2</sub>O, O<sub>3</sub>, CO<sub>2</sub>, O<sub>2</sub> and CH<sub>4</sub>. The concentration of N<sub>2</sub> is specified as well, EXPLANATION TO FOLLOW.

In the longwave, there are a much larger number of greenhouse gases that could be included, many of which have a very small individual impact. However, the purpose of CKDMIP is to evaluate the techniques used by schemes for generating CKD models based on the different requirements of weather and climate modelling, rather than to produce a single optimum CKD model that explicitly represents all the greenhouse gases that anyone might want to simulate. Therefore we have chosen to follow the pragmatic approach of Meinshausen et al. (2017). They stated that 94.5% of the anthropogenic greenhouse warming (in terms of radiative forcing) between 1750 and 2014 was due to increases in

**Table 1.** The four modelling applications of radiation schemes that we envisage would need to be targeted by a different CKD model. The present-day and variable trace-gas concentrations for these scenarios are provided in Table 2. Heating rates are evaluated at pressures down to the indicated ‘lowest pressure’, although note that the reference line-by-line calculations are performed down to lower pressures than these.

Application	Lowest pressure	Trace gas concs.
Limited-area NWP	4 hPa	Present-day (2020)
Global NWP	0.02 hPa	Present-day (2020)
Reanalysis	0.02 hPa	Variable
Climate	0.02 hPa	Variable

CO<sub>2</sub>, CH<sub>4</sub>, N<sub>2</sub>O, CFC-11 and CFC-12, with the remaining 5.5% being attributable to 38 further gases. Their ‘Option 2’ approximately represents the radiative forcing of these 38 gases by artificially increasing the concentration of CFC-11 (by around a factor of 3.9 in the present day); indeed, the CMIP6 (Coupled Model Intercomparison Project Phase 6) historic concentrations and future scenarios are available with these ‘CFC-11-equivalent’ concentrations. From Cycle 47R1, ECMWF’s Integrated Forecasting System will take this approach, using concentrations from the CMIP6 SSP3-7.0 scenario (O’Neill et al., 2016). For the CKDMIP activity we therefore limit ourselves to consideration of 9 gases: H<sub>2</sub>O, O<sub>3</sub>, CO<sub>2</sub>, O<sub>2</sub>, CH<sub>4</sub>, N<sub>2</sub>O, CFC-11, CFC-12 and N<sub>2</sub>. Note that the longwave absorption of O<sub>2</sub> and N<sub>2</sub> are included; their impacts on net TOA irradiance are around 0.11 and 0.17 W m<sup>-2</sup>, respectively (Höpfner et al., 2012).

### 2.2 Numerical weather prediction

Table 1 lists the four main applications for which we envisage that CKD models could be optimized. The first two correspond to present-day Numerical Weather Prediction (NWP) at the local and global scale. Both need to represent variable water vapour and ozone, but to a good approximation can assume all other gases to have a constant mole fraction, or to vary as a function of pressure alone. (Note that since the atmosphere is an ideal gas to a good approximation, we assume the mole fraction of a gas to be equal to its volume mixing ratio.) In principle, this allows the number of g-points to be reduced since, for example, all the well-mixed gases could be merged into a single ‘hybrid’ or ‘composite’ gas whose optical properties vary as a function of temperature and pressure alone (e.g., Ritter and Geleyn, 1992; Niemelä et al., 2001).

In terms of the present-day concentrations of the well-mixed gases, we assume that O<sub>2</sub> and N<sub>2</sub> have a constant volume mixing ratio of 0.20946 and 0.78102 mol mol<sup>-1</sup>, respectively, independent of pressure (Jones and Schoonover, 2002). These concentrations are also assumed for all past and future scenarios in section 2.3. The present-day surface concentrations of the five other well-mixed gases are shown in

Table 2, and were taken from the CMIP6 SSP3-7.0<sup>1</sup> scenario for calendar year 2020. The vertical profiles of these gases are discussed in section 3.2.

The difference between the two NWP applications listed in Table 1 is in the location of the model top. The model top quoted for all current configurations of the ECMWF model and all global configurations of the Met Office model used for weather and climate is 0.01 hPa (around 80 km). In the case of the ECMWF model this actually means that the highest model layer spans the pressure range 0–0.02 hPa. Since the temperature of the highest layer of a model is typically strongly affected by the ‘sponge’ (Shepherd et al., 1996), we limit evaluation of heating rates to pressures greater than 0.02 hPa. For the limited-area NWP application we evaluate heating rates only for pressures greater than 4 hPa, comparable to the model top used in the Met Office UKV model.

### 2.3 Reanalysis and climate modelling

The final two applications listed in Table 1 consider the five anthropogenic greenhouse gases to vary over the ranges shown in Table 2. The longest atmospheric reanalysis datasets have been generated back to the mid-19th century (e.g. Compo et al., 2011), so we propose that CKD models to be used for such an application should be able to simulate the radiative impacts of gas concentrations from 1850 (essentially the preindustrial concentrations of these gases) to 2030, i.e. the near-future at the time of writing (concentrations taken from the CMIP6 SSP3-7.0 scenario). The assessment of the CKD models for reanalysis would evaluate them for the ‘preindustrial’ and ‘present-day’ scenarios listed in Table 2, but also test the radiative impact of varying each of the gases in turn while keeping the others at present-day concentrations.

CKD models used for climate modelling should cover a much wider range. For CO<sub>2</sub> we match the range of 0.5–8 times preindustrial concentrations used in RFMIP (Pincus et al., 2016). For the other four well-mixed anthropogenic greenhouse gases we require the capability to simulate the minimum concentrations found in the last million years, which occurred at glacial maxima, up to the maximum concentrations found in any of the CMIP6 future scenarios, which extend until 2250. The resulting ranges are shown in the bottom row of Table 2. The top row also shows the individual scenario of a glacial maximum, with the values for CO<sub>2</sub> and CH<sub>4</sub> taken from Petit et al. (1999) and for N<sub>2</sub>O from the shorter period reported by Schilt et al. (2010). The fourth row shows a ‘future’ scenario consisting of worst-case conditions for 2110 by extracting the maximum concentrations from any of the CMIP6 scenarios at this time. In this year, the concentration of CH<sub>4</sub> peaks at 3500 ppbv in

the SSP3-7.0 scenario, and equivalent CFC-11 peaks at 2000 pptv in the SSP5-8.5 scenario.

## 3 Generating datasets

Four datasets are provided in CKDMIP, listed in Table 3. Each consists of profiles of layerwise spectral optical depth due to individual gases. The first two (Evaluation-1 and Evaluation-2) each consist of 50 realistic profiles of temperature, water vapour and ozone (described in section 3.1), accompanied by vertical profiles of the well-mixed gases (described in section 3.2. Evaluation-1 is provided to participants and may be used to train individual CKD models, while Evaluation-2 is held back to provide independent evaluation. Section 3.3 describes the last two datasets, which could also be useful in the training of new CKD models. Section 3.4 then describes how the profiles of spectral optical depth were computed for each dataset.

### 3.1 Temperature, humidity and ozone

For evaluating radiation schemes in RFMIP, Pincus et al. (2016) extracted a set of 100 contrasting atmospheric profiles from the 60-layer ERA-Interim reanalysis dataset, whose highest model level spans the pressure range 0–0.1 hPa. As well as being ten times higher than the pressure of the highest model level in the current ECMWF and Met Office global models, this vertical grid not sufficient to fully resolve the strong peak in atmospheric heating and cooling rates that occurs at the stratopause, nor to test solar absorption by molecular oxygen.

Therefore, we have selected a new set of temperature, pressure, humidity and ozone profiles from the 25,000 ‘NWP-SAF’ profiles of Eresmaa and McNally (2014), which they extracted from ECMWF operational model forecasts in 2013 and 2014. By this time the model used 137 layers with the highest layer spanning pressures 0–0.02 hPa, as in its current configuration. The 50 profiles of the ‘Evaluation-1’ dataset consist of 33 randomly taken from the 5,000-profile subset that Eresmaa and McNally (2014) themselves selected to maximize variations in temperature. An additional 17 profiles are selected to contain the extreme values (both maximum and minimum) in the entire dataset of (a) temperature in the layer nearest the surface, (b) temperature at 500 hPa, (c) temperature at 100 hPa, (d) temperature at 10 hPa, (e) temperature at 1 hPa, (f) specific humidity at 500 hPa, (g) specific humidity at 100-hPa (maximum only), (h) ozone concentration at 10-hPa, and (i) ozone concentration at 1 hPa.

It was apparent from inspection of the data that there was virtually no variability in stratospheric water vapour in the ECMWF model at the time the NWP-SAF profiles were generated, which is a problem for training and evaluating a gas-optics model. Therefore, additional variability has been

<sup>1</sup>‘SSP3-7.0’ is the ‘regional rivalry’ Shared Socioeconomic Pathway of CMIP6, with an anthropogenic radiative forcing of 7.0 W m<sup>-2</sup> in 2100.

**Table 2.** Surface volume mixing ratios of the five main anthropogenic greenhouse gases for various scenarios considered in CKDMIP, where ‘CFC-11 equivalent’ is an artificially increased CFC-11 concentration to represent 38 further greenhouse gases. The numbers in brackets indicate the multiple of preindustrial concentrations. The first four individual scenarios will be used to test CKD models developed for the four applications in Table 1: models for NWP will be tested on present-day concentrations, models for reanalysis will be tested also on preindustrial concentrations, and models for climate will be tested on all four. The final two lines indicate the range of concentrations over which CKD models generated for reanalysis and climate applications will be tested.

Scenario	CO <sub>2</sub> ppmv	CH <sub>4</sub> ppbv	N <sub>2</sub> O ppbv	CFC-11 equivalent pptv	CFC-12 pptv
<i>Individual scenarios</i>					
Glacial maximum	180 (0.64)	350 (0.5)	190 (0.7)	32	0
Preindustrial	280 (1)	700 (1)	270 (1)	32	0
Present-day (2020)	415	1921	332	861	495
Future (2110)	1120 (4)	3500 (5)	405 (1.5)	2000	200
<i>Concentration ranges covered by two of the applications</i>					
Reanalysis (1850–2030)	280–448 (1–1.6)	700–2100 (1–3)	270–340 (1–1.26)	0–950	0–550
Climate	140–2240 (0.5–8)	350–3500 (0.5–5)	190–540 (0.7–2)	0–2000	0–550

**Table 3.** The four spectral optical depth datasets generated as part of CKDMIP, where  $T$  is temperature,  $p$  is pressure and  $q$  is specific humidity.

Name	Purpose	Layers	$T$ profiles	Description
Evaluation-1	Training & evaluation	54	50	Realistic profiles selected from NWP-SAF dataset
Evaluation-2	Independent evaluation	54	50	Further profiles selected from NWP-SAF dataset
MMM	Training	52	3	Median, minimum and maximum of NWP-SAF $T$ , H <sub>2</sub> O & O <sub>3</sub> profiles
Idealized	Generating look-up tables	53	11	Idealized profiles regularly spaced in $T$ , $\log p$ and $\log q$

added by multiplying the humidity profiles by the following function of pressure,  $p$ :

$$f(p, r) = \exp \left[ r \times \frac{1 - \operatorname{erf} \left( \frac{p - 100 \text{ hPa}}{50 \text{ hPa}} \right)}{2} \right], \quad (1)$$

where  $r$  is a random number drawn from a Normal distribution with mean of zero and standard deviation 0.25, and is constant for each individual profile. This function adds around 25% variability in the stratosphere and mesosphere, but leaves the troposphere virtually unchanged. Unrealistically low humidities have been removed by setting the minimum specific humidity to  $10^{-7} \text{ kg kg}^{-1}$ .

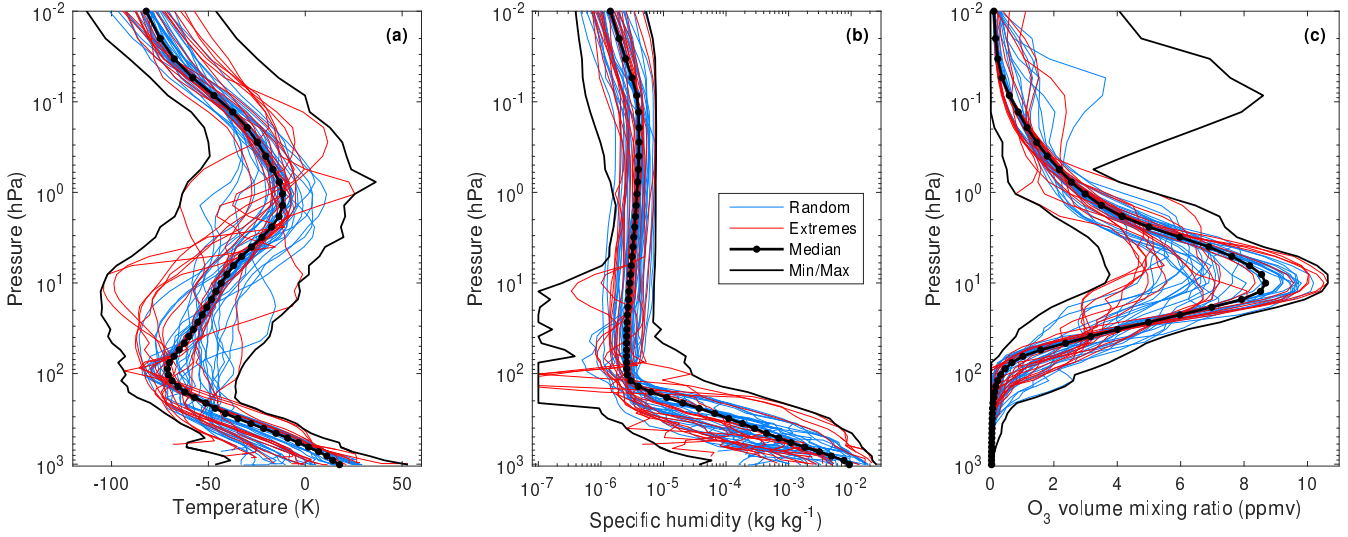
The resulting temperature, humidity and ozone mixing ratios are shown by the red and blue lines in Fig. 1. An additional ‘Evaluation-2’ dataset has been generated with a different set of 33 random profiles from the original 25,000, along with the 17 profiles containing second-most extreme values of the variables listed above. These will be used to provide independent evaluation of the CKD models.

Training and evaluating a CKD model is costly both in terms of computation and storage due to the high spectral resolution required, and 137 layers is more than needed for evaluating clear-sky radiative transfer. Therefore, we interpolate the profiles on to a coarser grid with 54 layers. We use the Line-By-Line Radiative Transfer Model (LBLRTM; Clough et al., 2005), version 12.8, which takes as input the temperature, pressure and gas concentrations at the interfaces

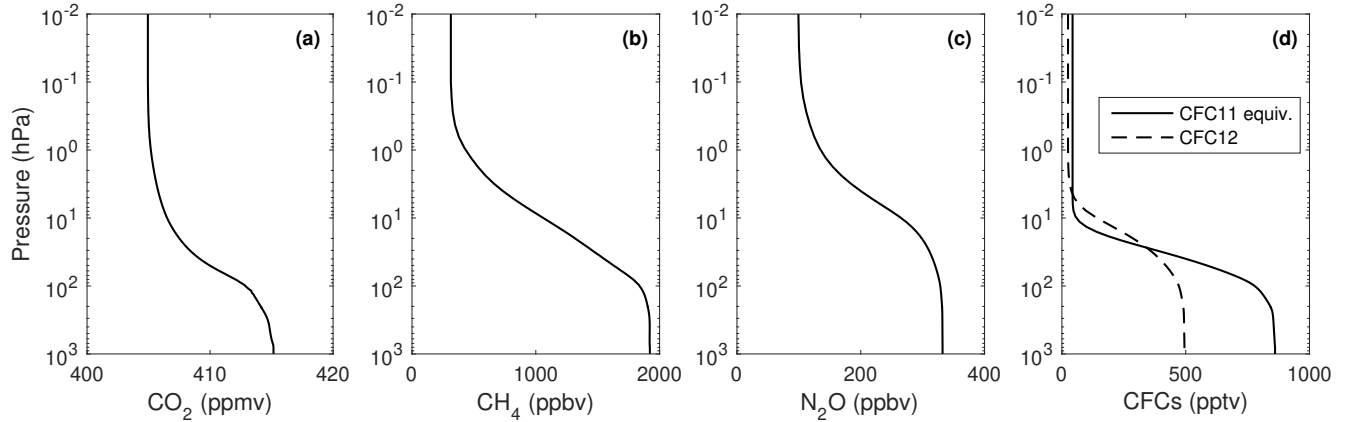
between layers. The highest two layers of the coarser grid are bounded by pressures of 0.0001, 0.01 and 0.02 hPa; the first of these represents the top-of-atmosphere since LBLRTM cannot compute gas properties at zero pressure. As shown in Table 1, 0.02 and 4 hPa are pressure surfaces marking the point at which evaluation of heating rates begins. We assign 15 layers between these two pressure surfaces, with the interfaces between them spaced linearly in  $p^{0.15}$  space, where  $p$  is pressure. The pressures defining the remaining layers vary according to the surface pressure  $p_s$ : we assign 35 layers between 4 hPa and the  $p_s/1.005$ , again spaced linearly in  $p^{0.15}$  space. Finally, a further two layers are added very close to the surface (bounded by  $p_s/1.005$ ,  $p_s/1.002$  and  $p_s$ ) in order to resolve sharp temperature gradients in the surface layer. The black dots in Fig. 1 mark the corresponding interfaces between layers for the median profiles described in section 3.3.

### 3.2 Well-mixed gases

Many weather and climate models assume a spatially constant mole fraction for each of the well-mixed gases, whereas for a little more realism they should decrease with height. The radiation scheme in the ECMWF model uses climatologies of these gases that vary with month, latitude and pressure, with the CO<sub>2</sub> and CH<sub>4</sub> climatologies taken from the MACC analysis system (Inness et al., 2013) and the N<sub>2</sub>O, CFC-11 and CFC-12 climatologies from the Cariolle chemistry model (Bechtold et al., 2009). Long-term changes due



**Figure 1.** Vertical profiles of the temperature, specific humidity and ozone concentrations for the ‘Evaluation-1’ dataset described in section 3.1.

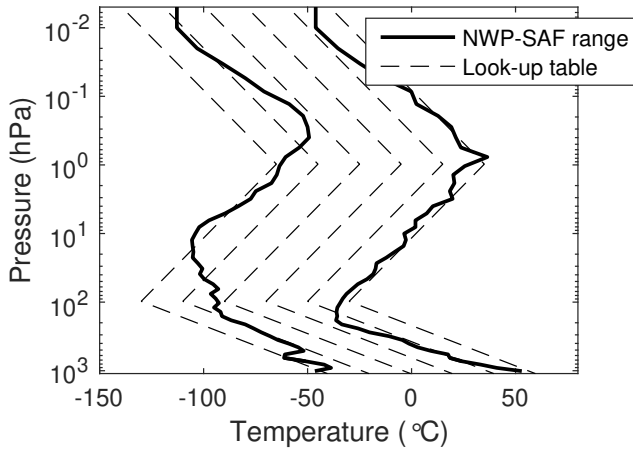


**Figure 2.** Vertical profiles of the five well-mixed greenhouse gases for the present-day (2020) surface concentrations listed in Table 2.

to anthropogenic emissions are represented by scaling these fields so that the global-mean surface values match either historic measurements (for hindcasts and reanalysis) or the CMIP6 SSP3-7.0 scenario (for operational forecasts from model cycle 47R1). We have averaged these climatologies globally and annually, and scaled them to the 2020 surface values in SSP3-7.0, to obtain the profiles shown in Fig. 2. Present-day CO<sub>2</sub> has a difference of 10 ppmv between the values at 1000 and 0.01 hPa. In the case of CFC-11 and CFC-12, the concentrations from the Cariolle model drop to almost zero in the upper stratosphere and mesosphere, which could be problematic for using them to train the pressure dependence in a CKD model. Therefore the profiles of these two gases have been artificially modified to fall to no less than 5% of their surface value. In order to obtain profiles with the surface concentrations shown in Table 2, we simply scale the profiles shown in Fig. 2.

### 3.3 Additional training datasets

Two additional datasets are shown at the bottom of Table 3, which are intended to facilitate the development of CKD schemes, while being consistent with the datasets that will be used to evaluate them. The ‘MMM’ dataset contains the optical properties of all nine gases but using the median, minimum and maximum temperature profiles derived from the entire 25,000-profile NWP-SAF dataset; these temperatures are shown by the black lines in Fig. 1. In the case of H<sub>2</sub>O and O<sub>3</sub> only, three concentration profiles are used for each temperature, corresponding also to the median, minimum and maximum of the NWP-SAF profiles (shown in Figs. 1b and 1c). For all other gases the present-day concentrations shown in Fig. 2 are used. The vertical grid is the same as for the Evaluation-1 and Evaluation-2 datasets, except that surface pressure is set to mean sea level pressure ( $p_s = 1013.25$  hPa),



**Figure 3.** The solid lines show the minimum and maximum temperatures of the NWP-SAF dataset, also shown in Fig. 1a. The dashed lines show the 6 idealized temperature profiles, 20 K apart, used in the ‘Idealized’ dataset in Table 3.

and the two layers very close to the surface are not used so that the total number of layers is 52 rather than 54.

The final ‘Idealized’ dataset contains absorption spectra for idealized temperature and concentration profiles that are intended to cover the full range of likely temperature, pressure and concentrations found in the atmospheres that any CKD model would be applied to. Therefore they can be used to populate look-up tables of molar absorption to be used by CKD models. We envisage that the maximum layer-mean pressure that needs to be accommodated by a radiation scheme is 1100 hPa, so construct a logarithmically spaced pressure profile of 53 elements, containing ten points per decade with a maximum layer-mean pressure of 1100 hPa. At each pressure, 6 temperatures are simulated spanning a 100 K range at 20 K intervals. We use idealized temperature profiles shown in Fig. 3 that are intended to encompass the maximum and minimum temperatures found in the NWP-SAF dataset. For all gases, absorption spectra are computed for mole-fraction profiles that are constant with pressure, using the present-day values for the five well-mixed gases shown in Table 2, and 5 ppmv for  $O_3$ . Since the molar absorption of these gases is very close to constant with concentration (see section 3.4), only one concentration needs to be simulated for each. In the case of water vapour whose absorption varies with concentration, we simulate 12 logarithmically spaced specific humidities from  $10^{-7}$  to  $10^{-1.5}$   $kg\ kg^{-1}$ , i.e. using two values per decade.

### 3.4 Line-by-line modelling

#### ADD DETAILS OF LBLRTM

An important practical consideration is to determine at what spectral resolution to produce the absorption spectra. They need to be fine enough resolution that the most narrow spectral lines are resolved and the resulting irradiance

and heating-rates profiles are an accurate benchmark, but also a manageable data volume for storage, processing and distribution. LBLRTM can inform the user of the spectral resolution it needs to resolve the lines at a particular pressure, and for  $CO_2$  at 0.01 hPa (the most important gas at the pressure where the lines are finest), it recommends a wavenumber resolution of 0.000036, 0.000156, 0.000292 and 0.000485  $cm^{-1}$  in the wavenumber ranges 0–350, 350–1300, 1300–1700 and 1700–3260  $cm^{-1}$ , respectively. This implies that more than 20 million spectral points are required. Using this resolution as a reference, we have experimented with degrading the spectral resolution in each of these four wavenumber ranges. Computing the heating rate error for each range, we find that the 350–1300  $cm^{-1}$  range is the only one of the four where anything like the LBLRTM-recommended resolution is required. Therefore, we have adopted spectral resolutions of 0.0002, 0.001 and 0.005  $cm^{-1}$  in three spectral ranges 0–1300, 1300–1700 and 1700–3260  $cm^{-1}$ . This leads to heating-rate errors of no more than around 0.005  $K\ d^{-1}$  (all of which occur in the upper stratosphere and mesosphere) in any of the four original wavenumber ranges, even for the most challenging scenario of 8 times preindustrial concentrations of  $CO_2$ . This leads to 7,211,999 spectral points in the longwave.

#### SHORTWAVE?

The TOA solar irradiance spectrum is provided at the same spectral resolution as the shortwave gas absorption spectra, extracted from the dataset of Coddington et al. (2016).

A further significant reduction in data volume is possible if the absorption cross-section per molecule is independent of the concentration of that gas, so varies only as a function of temperature and pressure. In this case, for well-mixed gases, the profile of layer-wise optical depth need only be provided for a single concentration profile; if optical depths are required for concentration profiles scaled by a constant, then the optical depths themselves may simply be scaled. We have computed absorption spectra for each gas over the full range of concentrations required in Table 2, and found that to a very good approximation, molar absorption can be treated as independent of concentration for all gases except water vapour. Therefore, for the well-mixed gases, absorption spectra are provided only for present-day concentrations. The CKDMIP software accordingly allows the user to scale the optical depth of each gas before performing radiative transfer calculations on the mixture.

The water vapour spectra have been computed assuming the widely used Mlawer-Tobin-Clough-Kneizys-Davies (MT\_CKD) continuum model (Mlawer et al., 2012), version 2.5. However, there is still considerable uncertainty on the strength of the water vapour continuum, particularly in the near infrared (Shine et al., 2016), and indeed this could be a source of difference between individual gas optics schemes and the reference calculations produced in CKDMIP. Therefore, for each dataset, we produce an additional set of water vapour files but with no representation of the continuum. If

needed, evaluation can be carried out using only the contribution from spectral lines, or alternatively different models of the continuum can be tried.

## 4 Radiative transfer assessment

### 4.1 Radiative transfer protocol

The CKDMIP software takes as input the spectral optical depths of each of a number of gases, optionally scales the optical depths of the well-mixed gases if a different concentration is required, and computes clear-sky aerosol-free irradiances (broadband or spectral) at layer interfaces for each of the test profiles. These can be used to compute broadband or spectral heating-rate profiles. The intention is that the radiative transfer equations are then the same as those used by large-scale atmospheric models, and the same solver will be used with the various CKD models in order that any differences to the line-by-line broadband irradiances are due to the representation of gas optics, not the details of the solver.

In the longwave we use a no-scattering solver with the following properties:

- Surface emissivity is assumed to be unity.
- The skin temperature of the surface is assumed to be equal to the air temperature at the base of the lowest atmospheric layer.
- Local thermodynamic equilibrium is assumed.
- The propagation of longwave radiation assumes a diffusivity factor of 1.66, i.e. radiation is assumed to propagate with discrete zenith angles of  $\pm 53^\circ$ .
- The temperature at layer interfaces is taken as input and a linear-in-optical-depth variation of the Planck function within each layer is assumed, leading to the use of Eqs. 6–12 of Hogan and Bozzo (2018).

The shortwave scheme has the following characteristics:

- The surface is assumed to be a Lambertian reflector with an albedo of 0.06.
- It uses a direct-beam calculation plus a two-stream diffuse calculation, with the Zdunkowski et al. (1980) coefficients characterizing the rate of exchange of energy between the three streams, and the Meador and Weaver (1980) solutions to the two-stream equations in individual layers.
- Calculations are performed at a number of solar zenith angles  $\theta_0$ , evenly spaced in  $\cos \theta_0$  space, thereby matching the distribution of solar zenith angles striking the Earth. Note that we do not account for the fact that individual test profiles at a particular latitude would each experience a different SZA distribution.
- No account is made for Earth curvature.

## 4.2 Error metrics

## 5 Results with RRTMG (?)

## 6 Conclusions

### Acknowledgments

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