



Funded by the  
European Union

Co-ordinated by  **ECMWF**

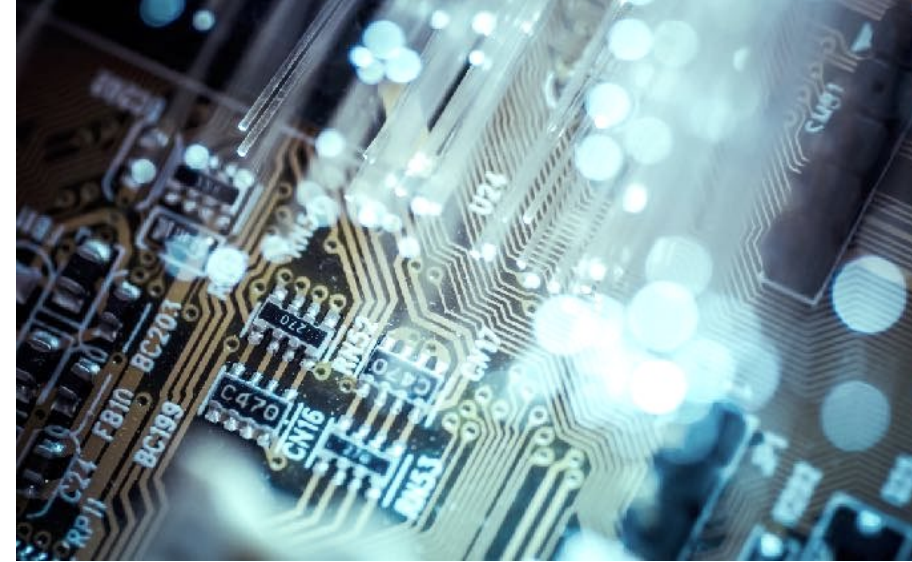
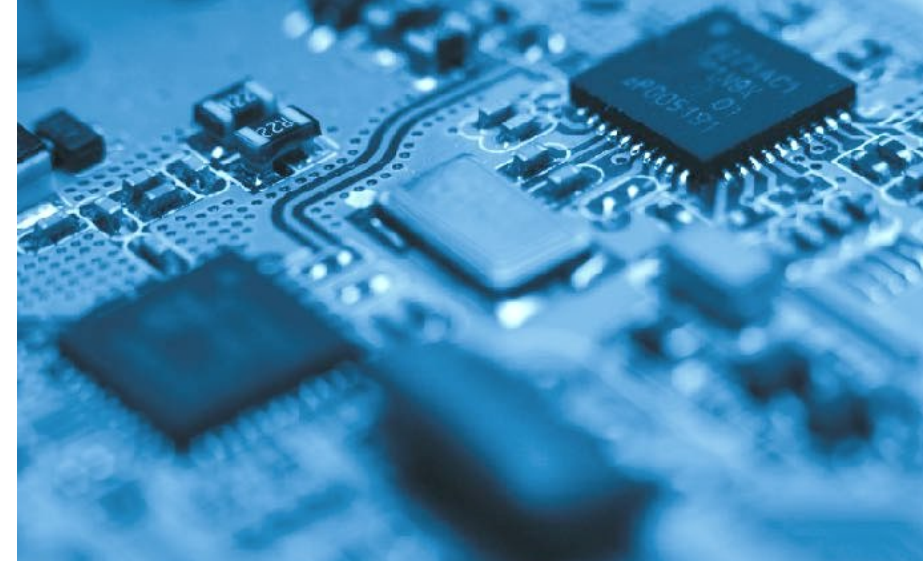
# ESCAPE



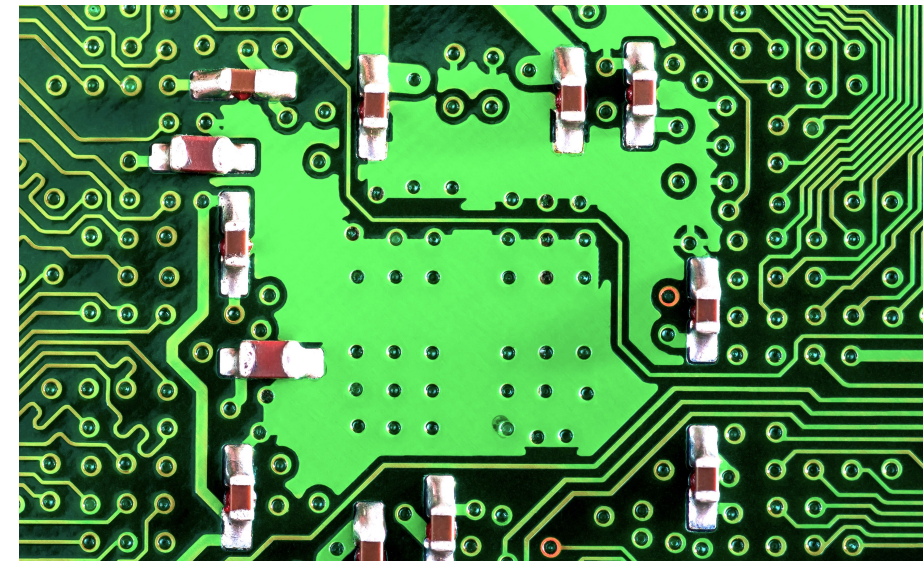


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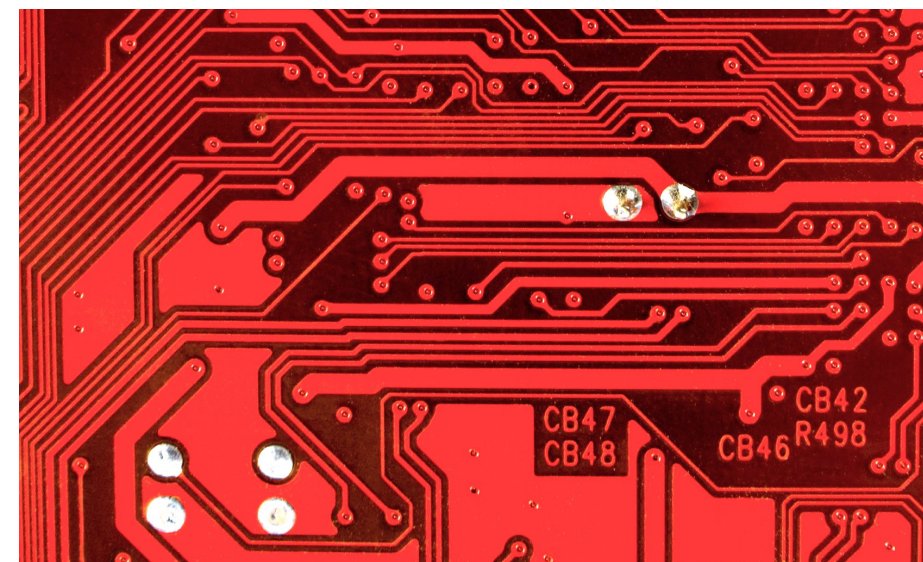


# ESCAPE 2



## Spectral Transform

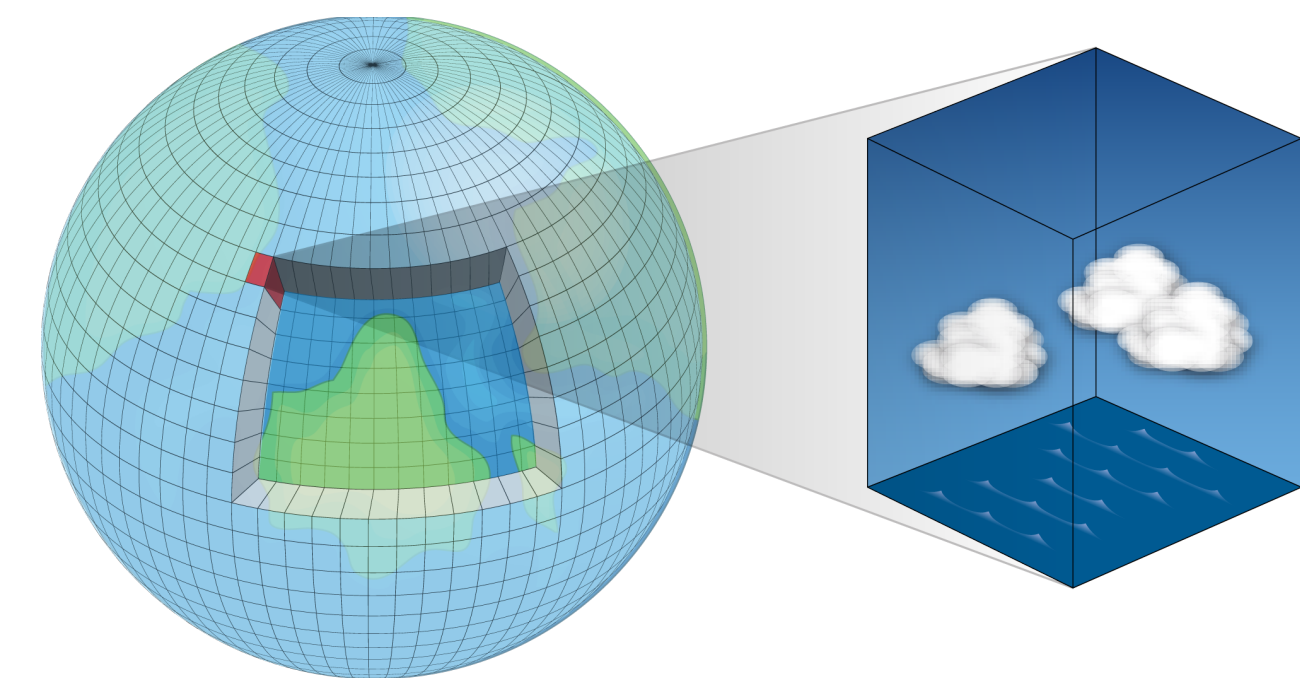
Andreas Mueller



# ESCAPE: Energy-efficient Scalable Algorithms for Weather Prediction at Exascale



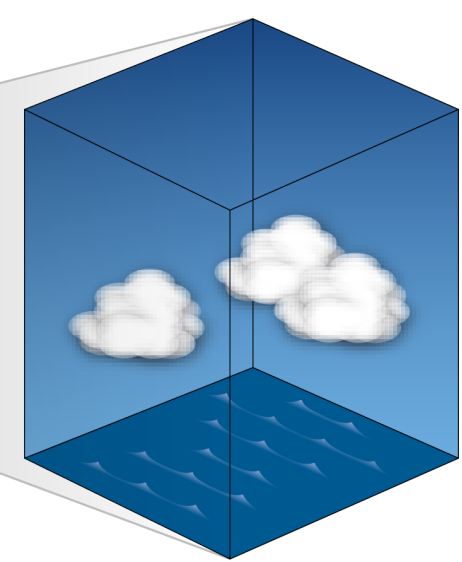
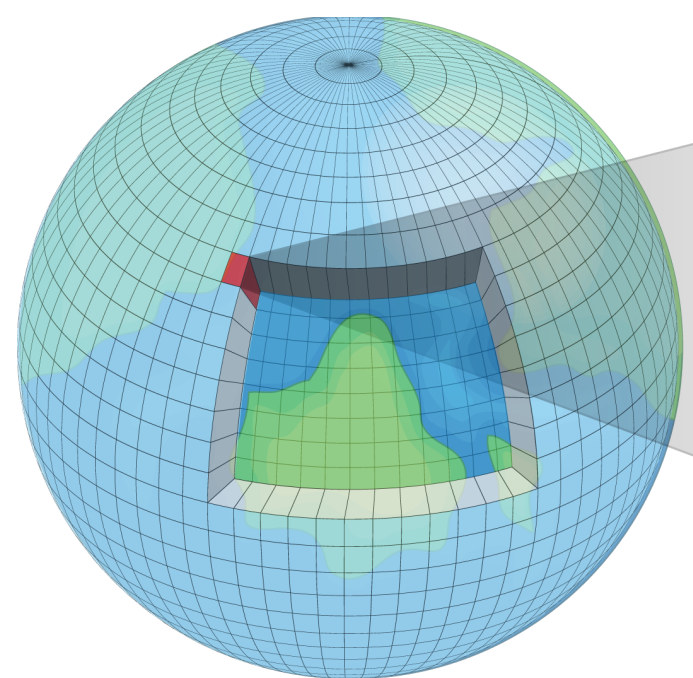
# ESCAPE: Energy-efficient Scalable Algorithms for Weather Prediction at Exascale



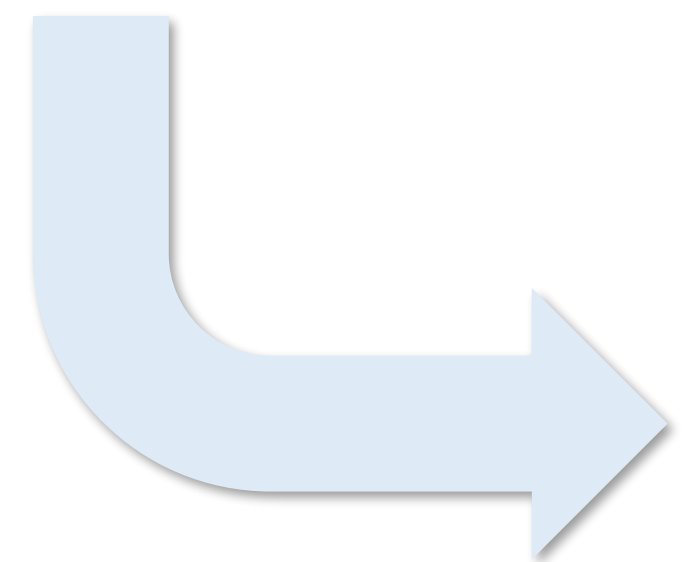
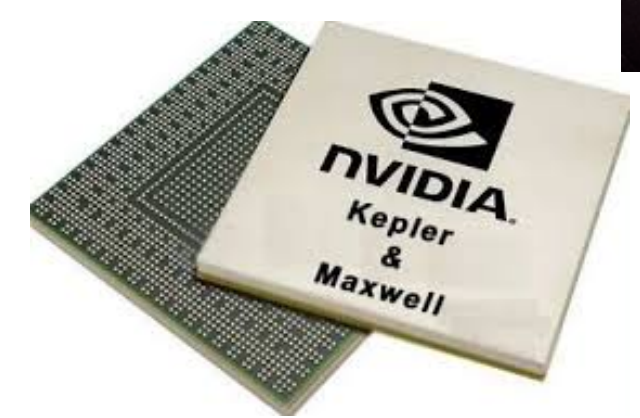
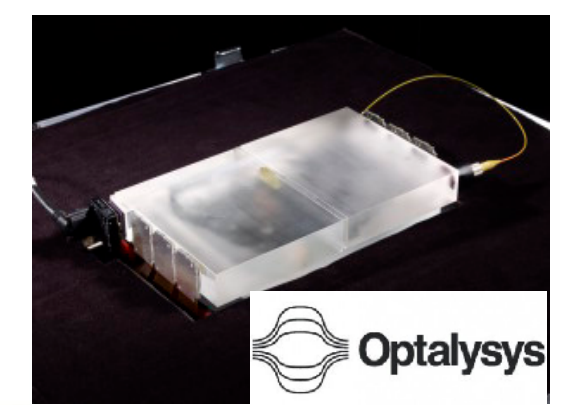
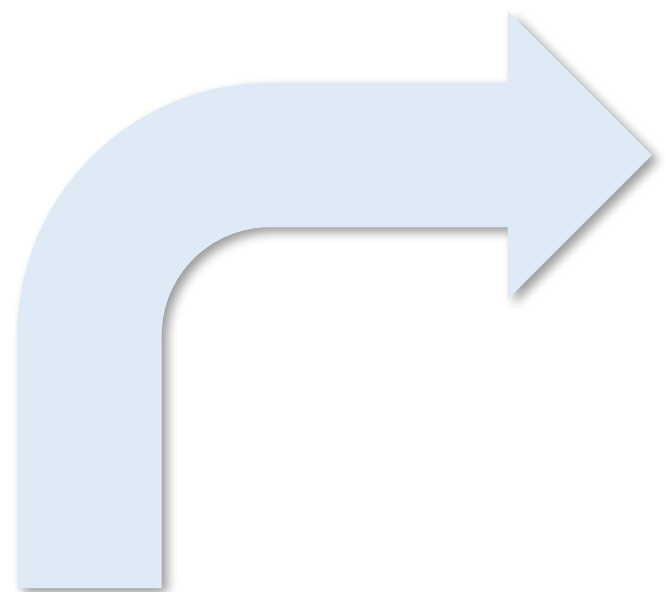
Extract model dwarfs...



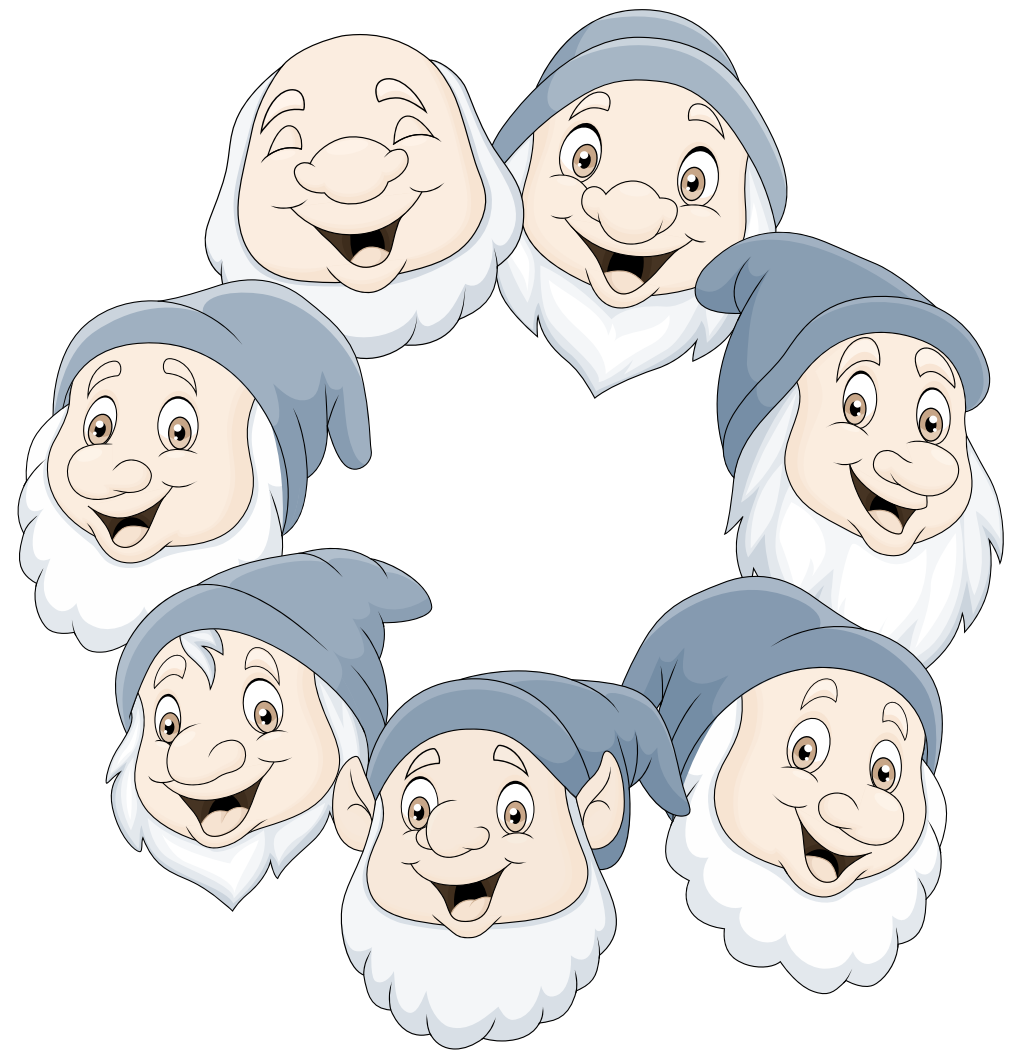
# ESCAPE: Energy-efficient Scalable Algorithms for Weather Prediction at Exascale



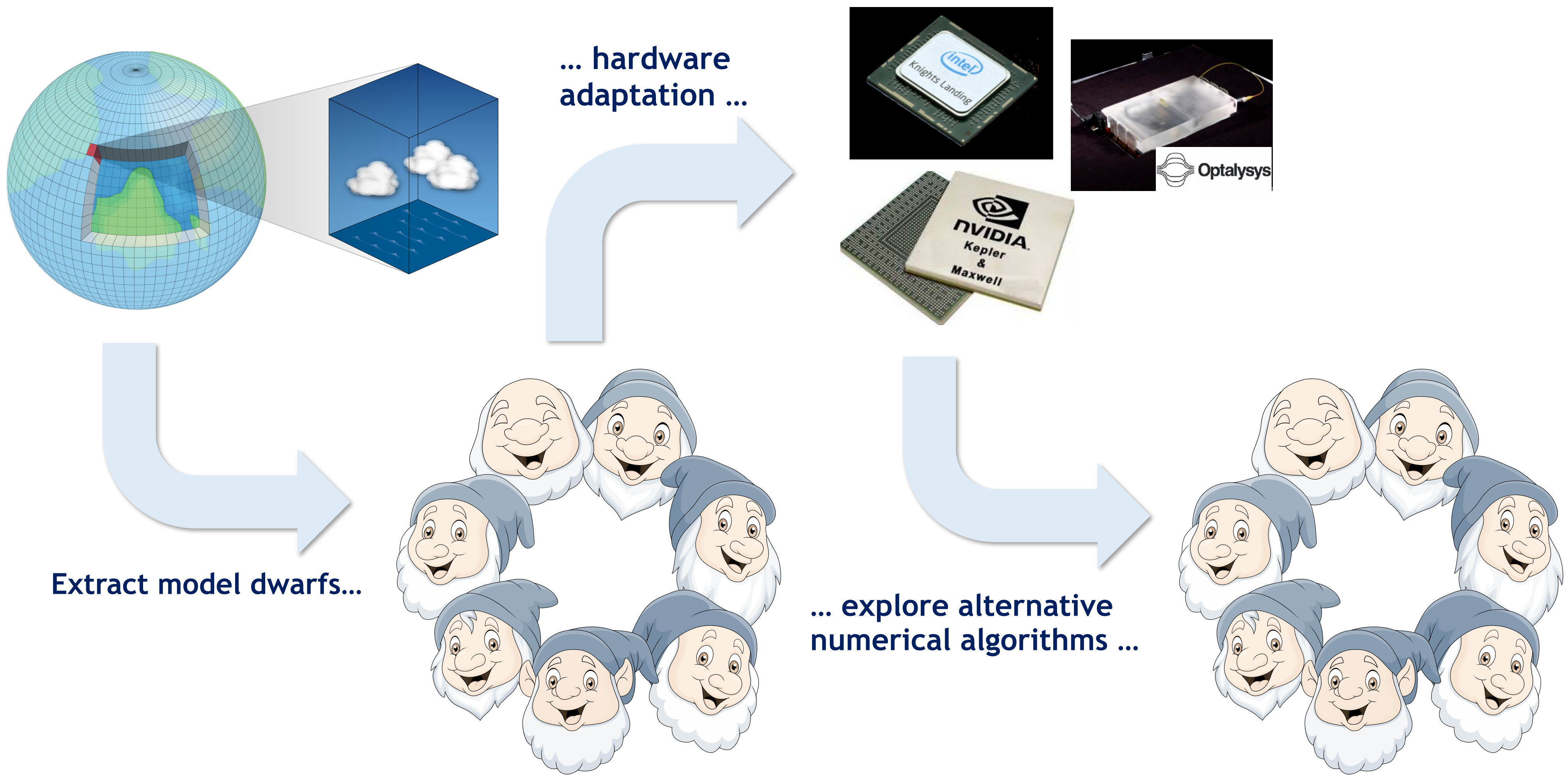
... hardware adaptation ...



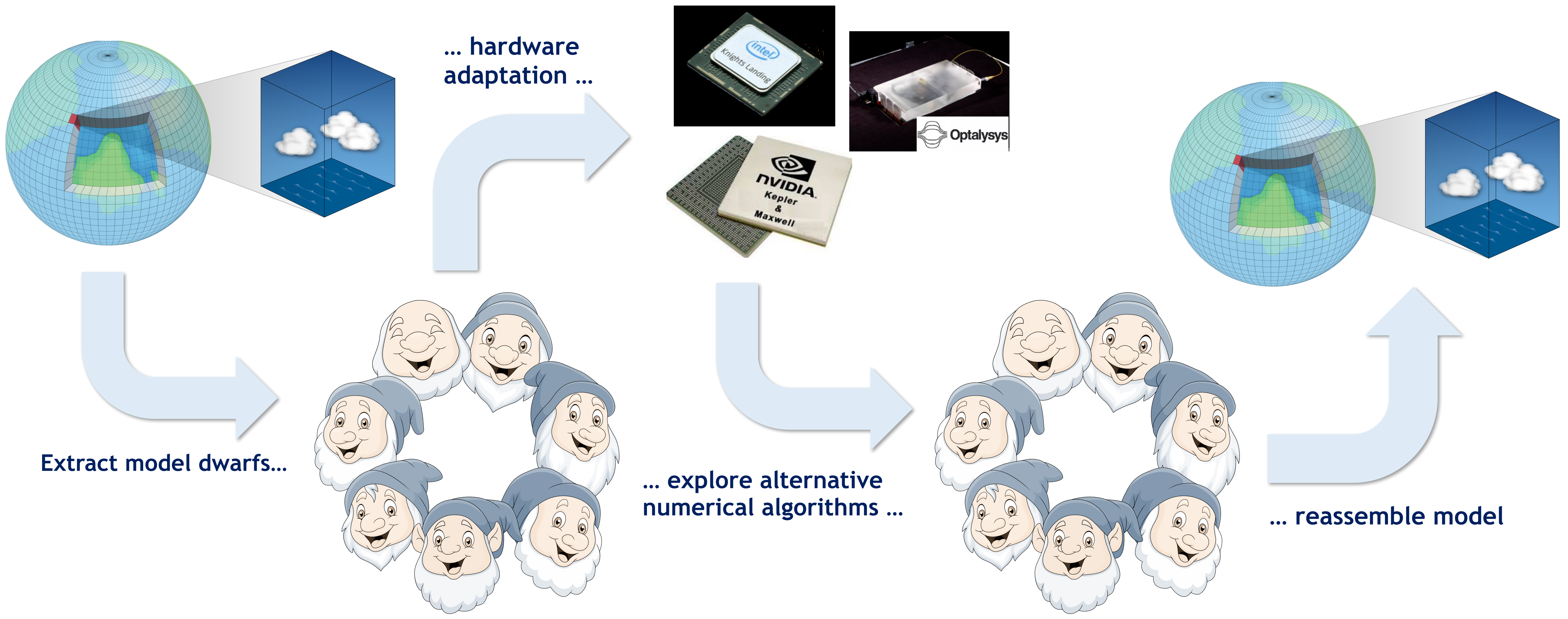
Extract model dwarfs...



# ESCAPE: Energy-efficient Scalable Algorithms for Weather Prediction at Exascale



# ESCAPE: Energy-efficient Scalable Algorithms for Weather Prediction at Exascale





# Overview

10 minutes

- Fourier transform
- Spectral transform

40 minutes

hands-on exercises:

- interactive web-app
- python notebook



# IFS (Integrated Forecast System)



technology applied at ECMWF

for the last 30 years

- spectral transform
- semi-Lagrangian
- semi-implicit



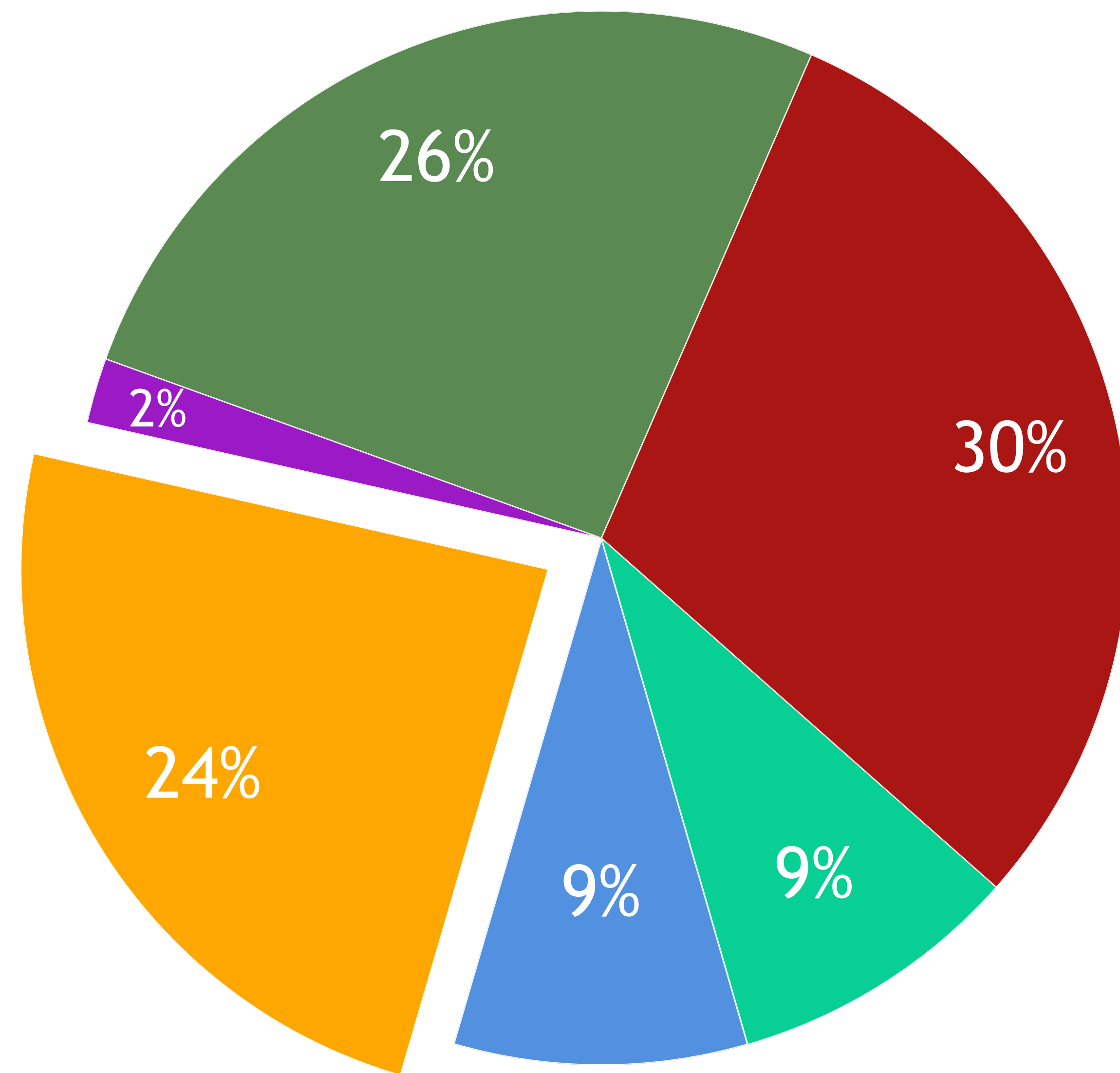
# IFS (Integrated Forecast System)

technology applied at ECMWF for the last 30 years

- spectral transform
- semi-Lagrangian
- semi-implicit

pie chart: % of runtime in 9km operational forecast

- spectral transform
- grid point dynamics
- wave model
- semi-implicit solver
- physics+radiation
- ocean model





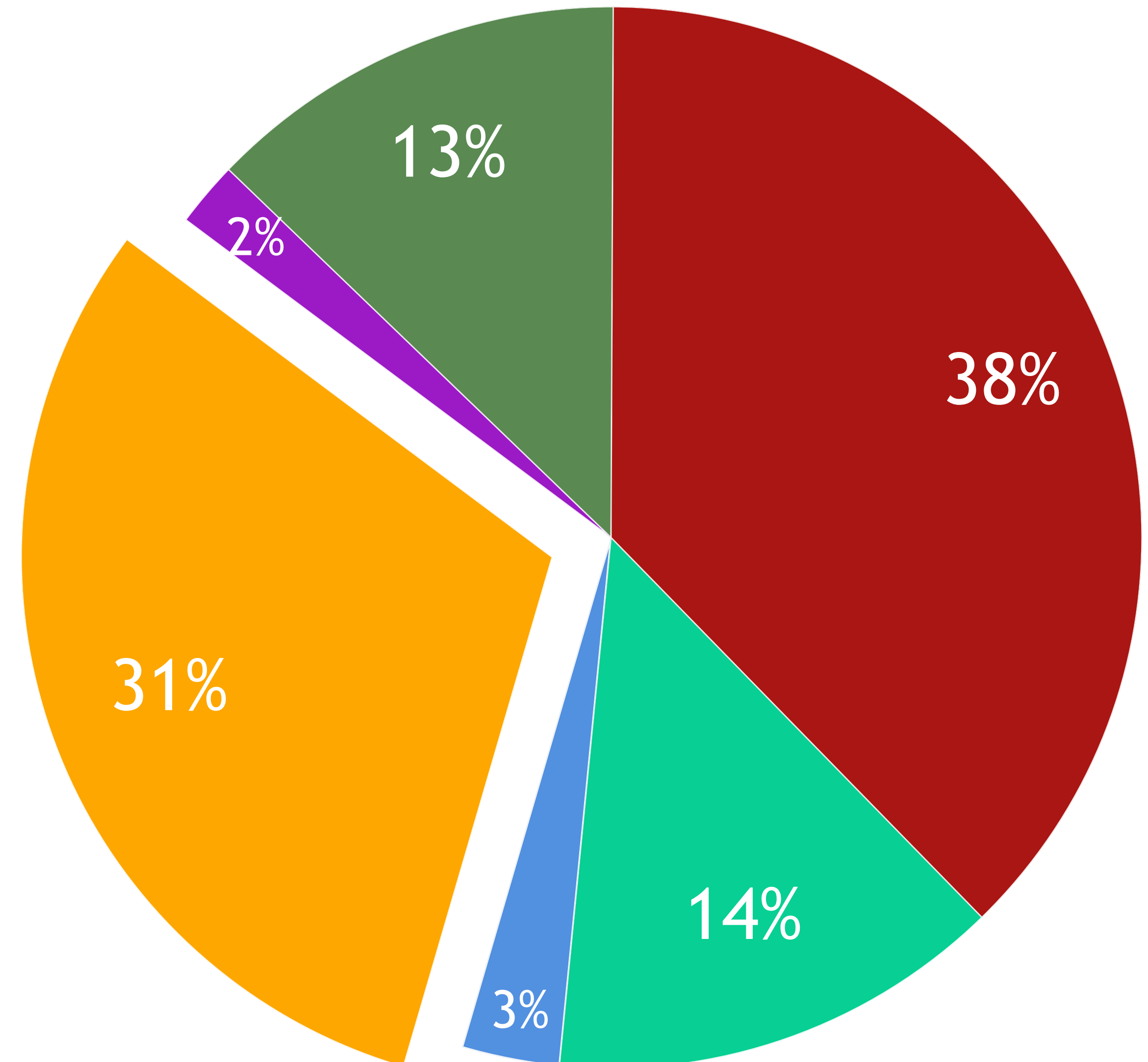
# IFS (Integrated Forecast System)

technology applied at ECMWF for the last 30 years

- spectral transform
- semi-Lagrangian
- semi-implicit

pie chart: % of runtime in 5km forecast (future operational)

- spectral transform
- grid point dynamics
- wave model
- semi-implicit solver
- physics+radiation
- ocean model





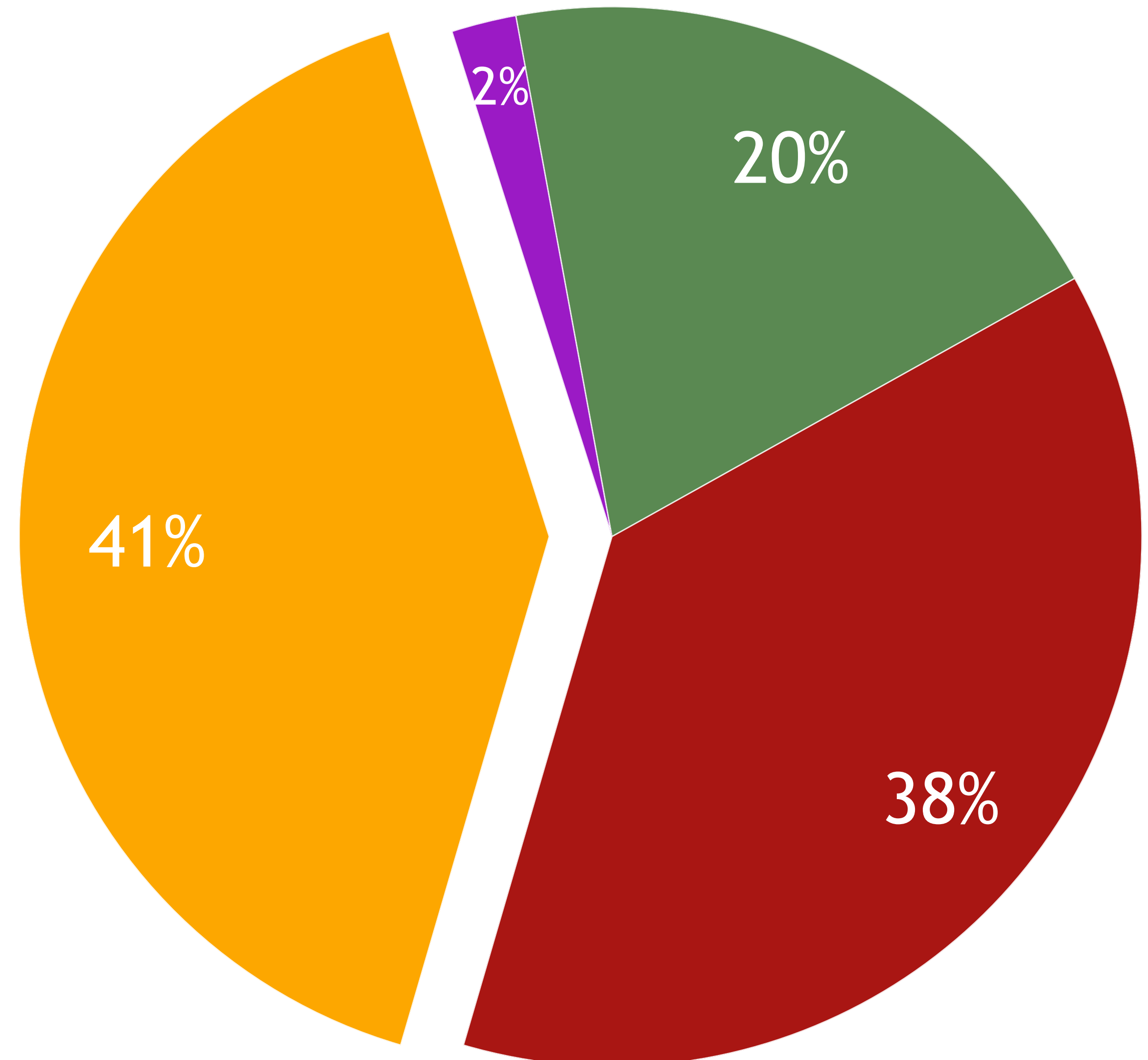
# IFS (Integrated Forecast System)

technology applied at ECMWF for the last 30 years

- spectral transform
- semi-Lagrangian
- semi-implicit

pie chart: % of runtime in 1.25km forecast (experiment, no ocean)

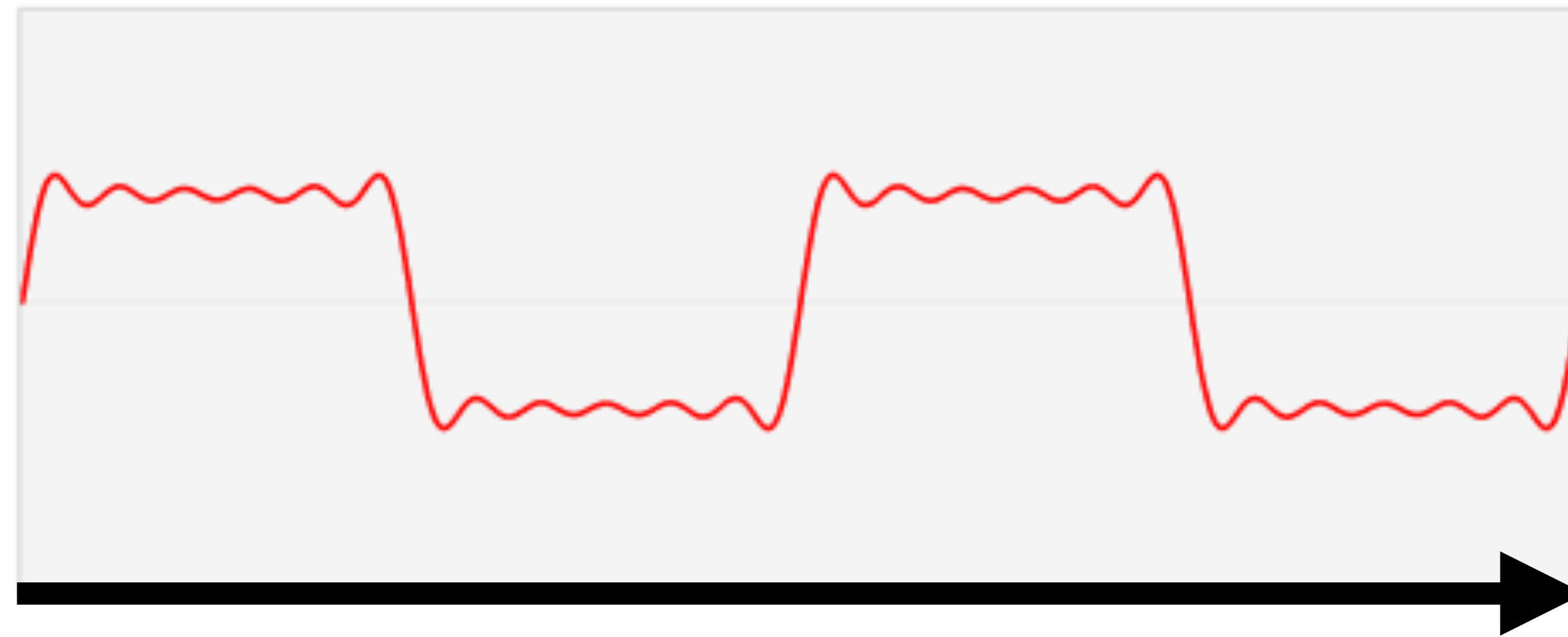
- spectral transform
- grid point dynamics
- wave model
- semi-implicit solver
- physics+radiation
- ocean model





# Fourier transform

Fourier transform = Spectral transform in 1D

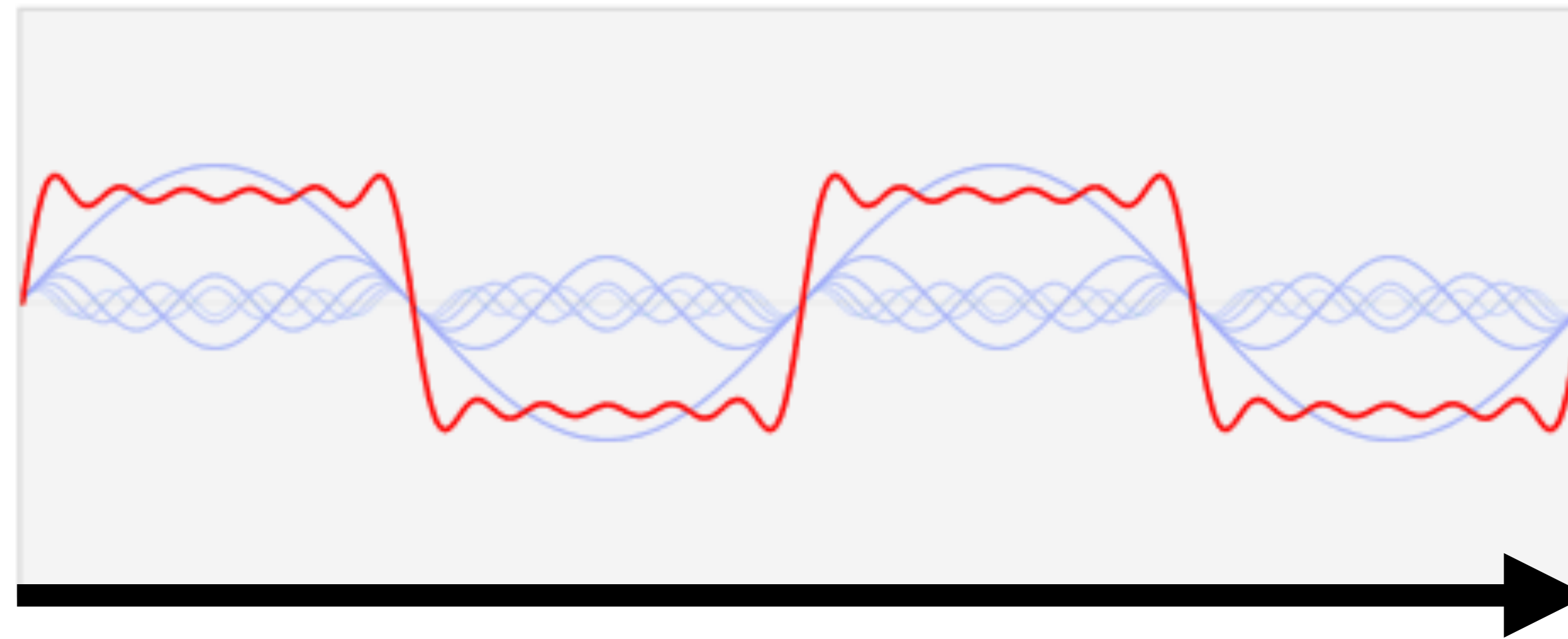


location x



# Fourier transform

Fourier transform = Spectral transform in 1D

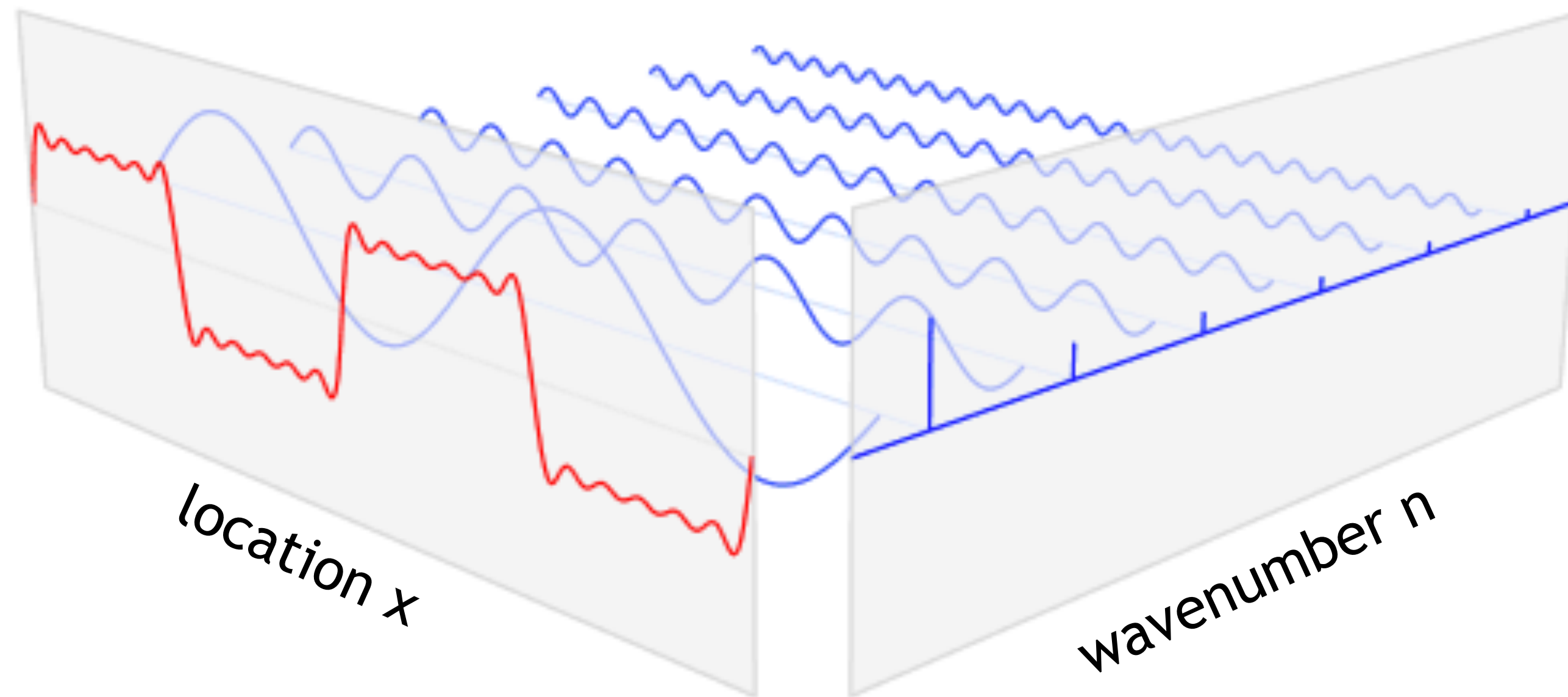


location x



# Fourier transform

Fourier transform = Spectral transform in 1D

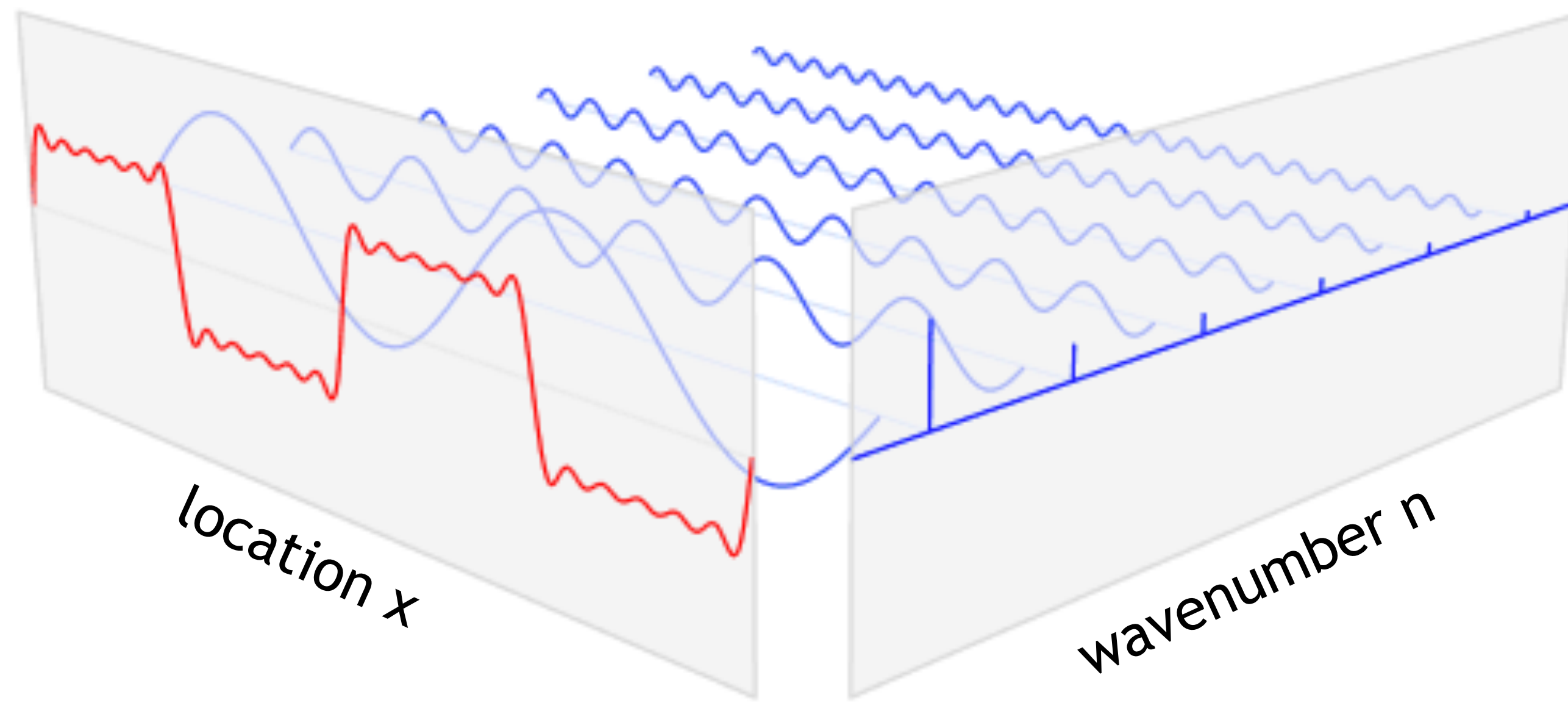


**grid point space**

**Fourier space**



# Fourier transform



function in grid  
point space

$$f(x) = \sum_n f_n \cdot e^{-2\pi i n x}$$

Fourier  
coefficients





# Fourier transform

function in grid  
point space

$$f(x) = \sum_n f_n \cdot e^{-2\pi i n x}$$

Fourier  
coefficients

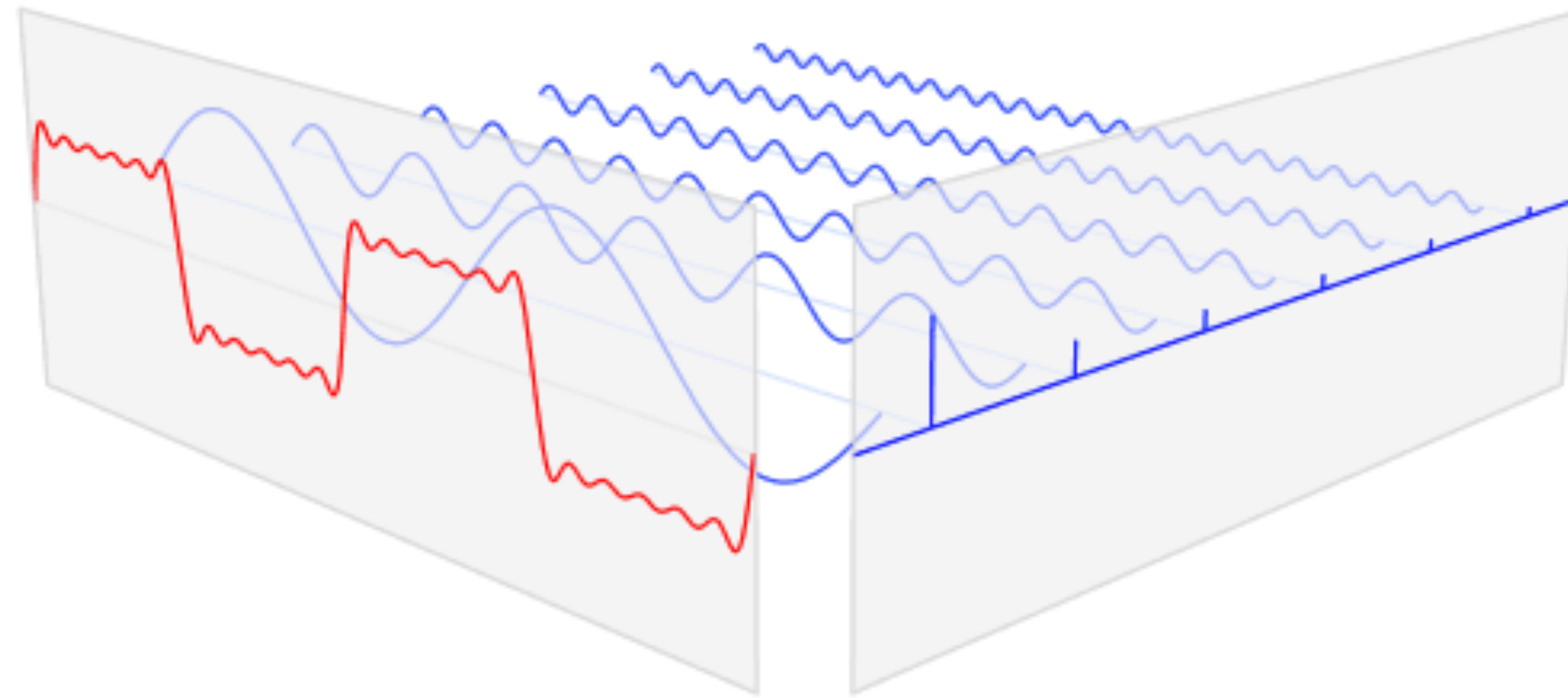
differentiation

$$\frac{df(x)}{dx} = \sum_n (-2\pi i n f_n) \cdot e^{-2\pi i n x}$$

simple  
multiplication

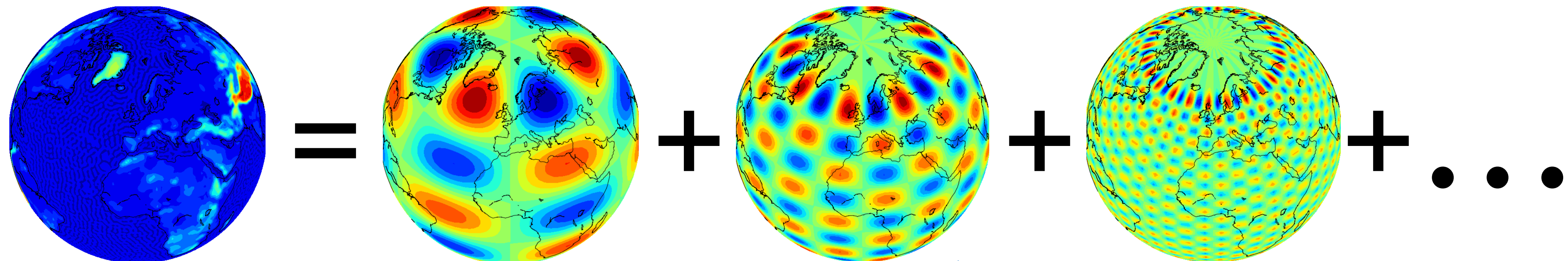


# on the sphere: spectral transform



grid point space

spectral space



spherical harmonics



# on the sphere: spectral transform

Spectral coefficients

Latitude Longitude

Grid point variable

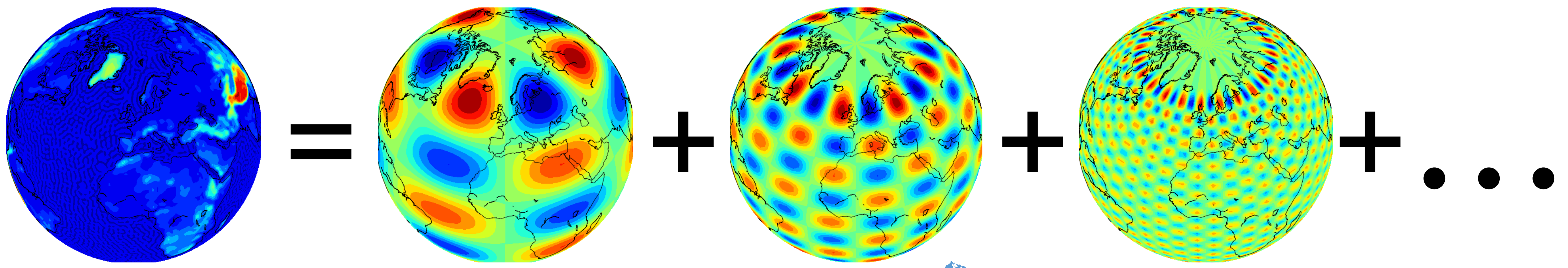
Spherical harmonics

$$f(\phi, \lambda) = \Re \left( \sum_{m=0}^M \sum_{n=m}^M f_{m,n} Y_n^m(\phi, \lambda) \right)$$

m: zonal wavenumber  
 n: total wavenumber  
 M: truncation

grid point space

spectral space



spherical harmonics



# on the sphere: spectral transform

Spectral coefficients

Grid point variable    Latitude    Longitude    Spherical harmonics

$$f(\phi, \lambda) = \Re \left( \sum_{m=0}^M \sum_{n=m}^M f_{m,n} Y_n^m(\phi, \lambda) \right)$$

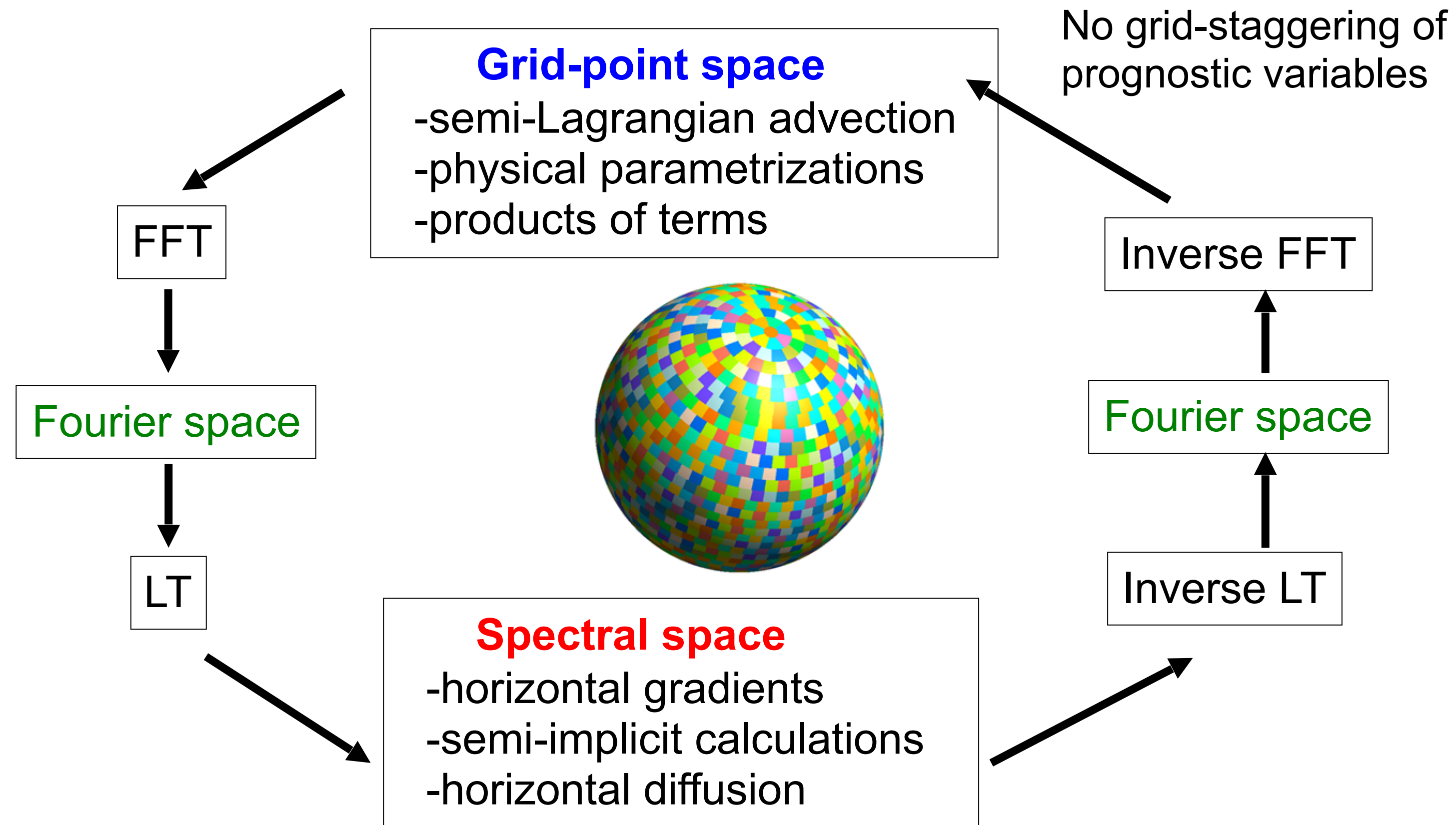
m: zonal wavenumber  
n: total wavenumber  
M: truncation

Legendre polynomials

$$f(\phi, \lambda) = \Re \left( \underbrace{\sum_{m=0}^M e^{im\lambda}}_{\text{Fourier transform}} \underbrace{\sum_{n=m}^M f_{m,n} P_n^m(\phi)}_{\text{Legendre transform}} \right)$$



# time step in IFS



FFT: Fast Fourier Transform, LT: Legendre Transform



# hands-on session

**for everyone: interactive web-app about spectral transform**

open in a browser: [anmrde.github.io/spectral](https://anmrde.github.io/spectral)

**optional: step 2: Python course**

in the classroom:

`/home/users/swx18100/Monday_training/spectral/install.sh`

in the cloud:

<https://notebooks.azure.com/anmrde/libraries/tcnm2019>

click on clone

**files:**

`exercises.ipynb`, `TCNM2019.ipynb`: Python notebook with exercises

`solution.ipynb`, `TCNM2019solution.ipynb`: notebook including sample solutions



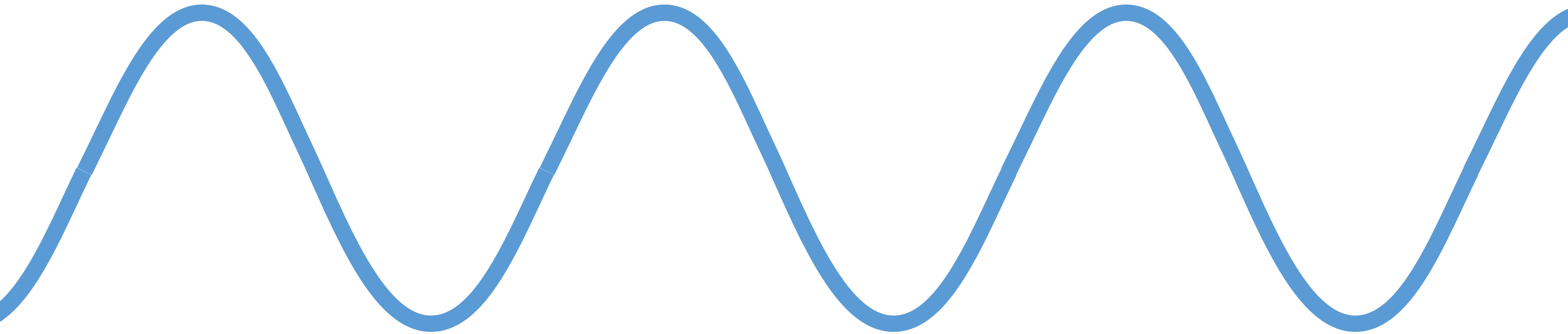
# aliasing

**Issue:** multiplication of two variables produces shorter waves than grid can handle



# aliasing

wave generated in spectral space



**Issue:** multiplication of two variables produces shorter waves than grid can handle

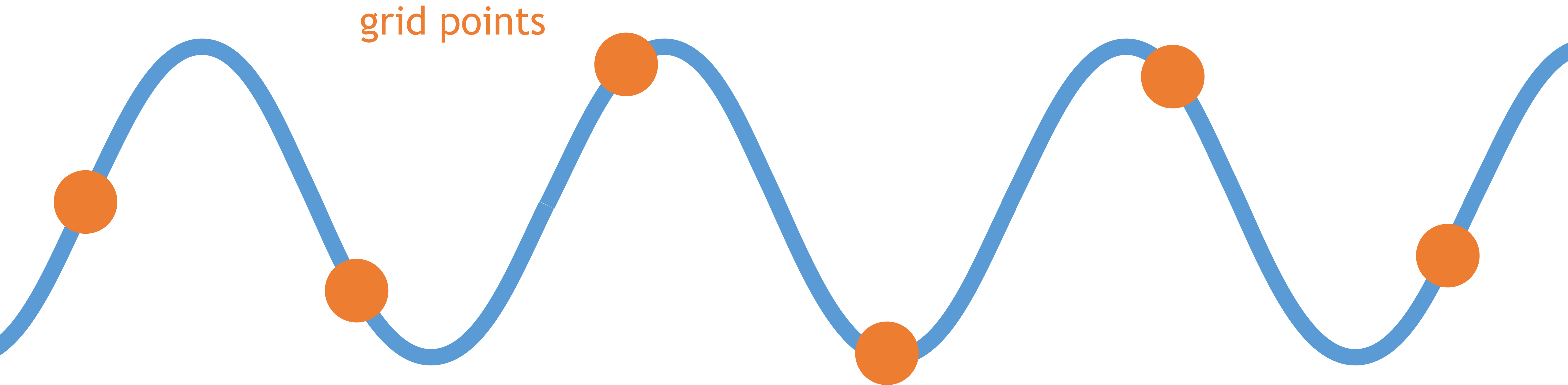




# aliasing

wave generated in spectral space

grid points



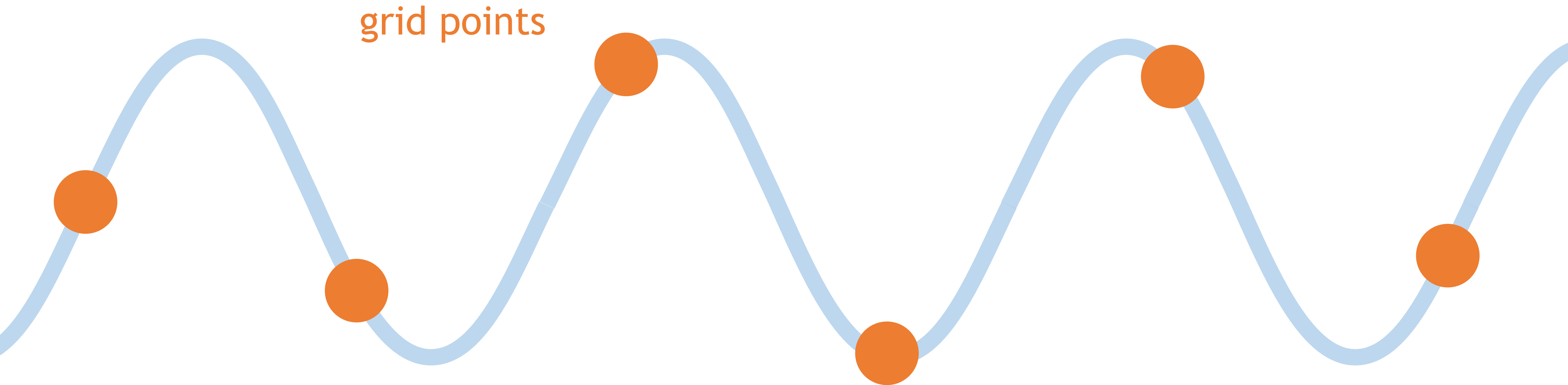
**Issue:** multiplication of two variables produces shorter waves than grid can handle



# aliasing

wave generated in spectral space

grid points



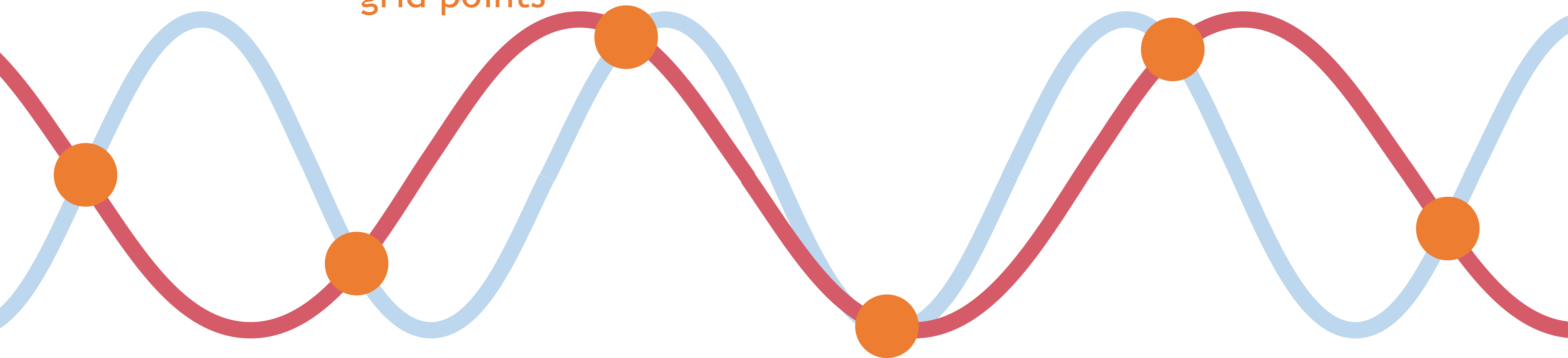
**Issue:** multiplication of two variables produces shorter waves than grid can handle



# aliasing

wave generated in spectral space

grid points



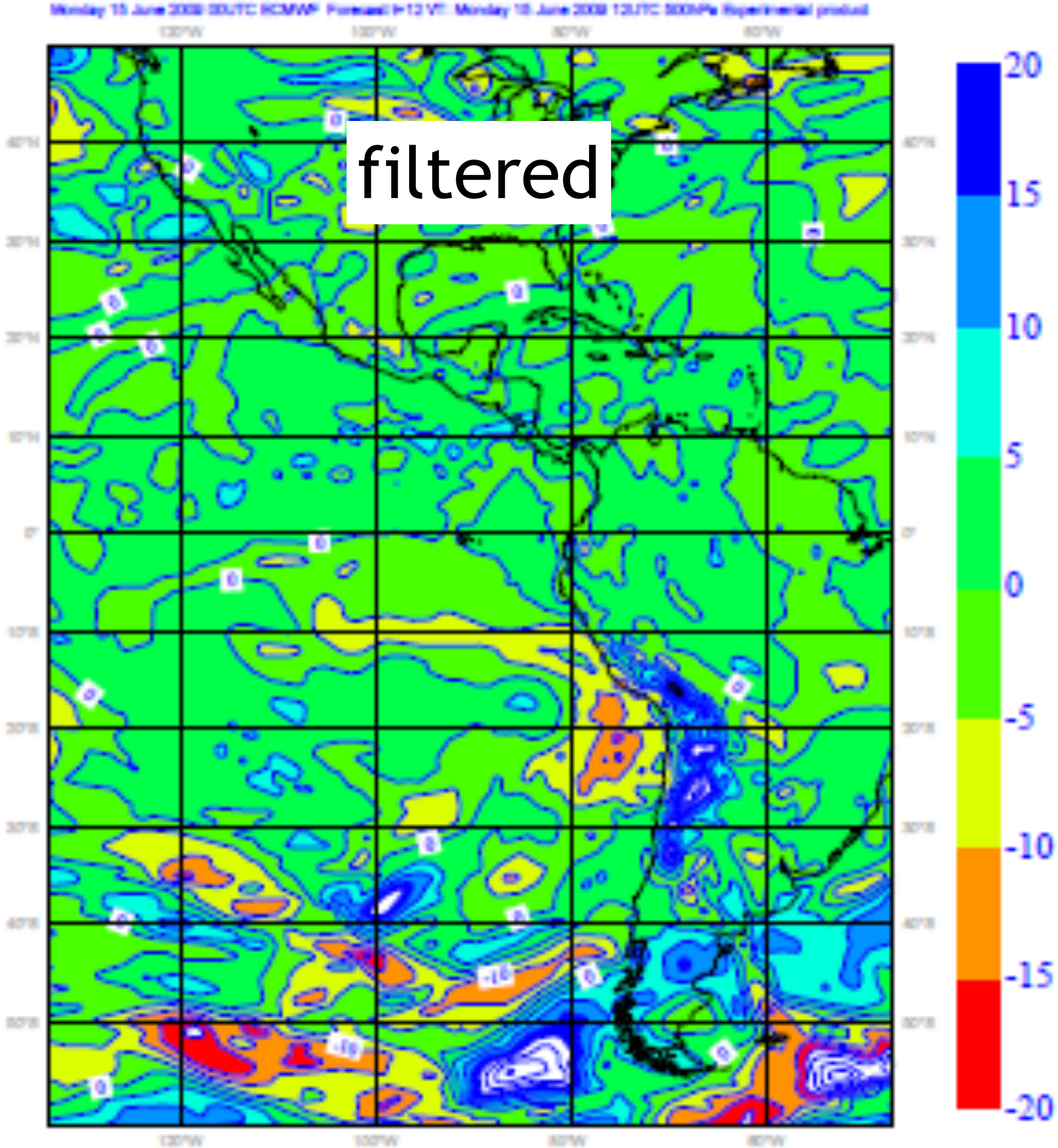
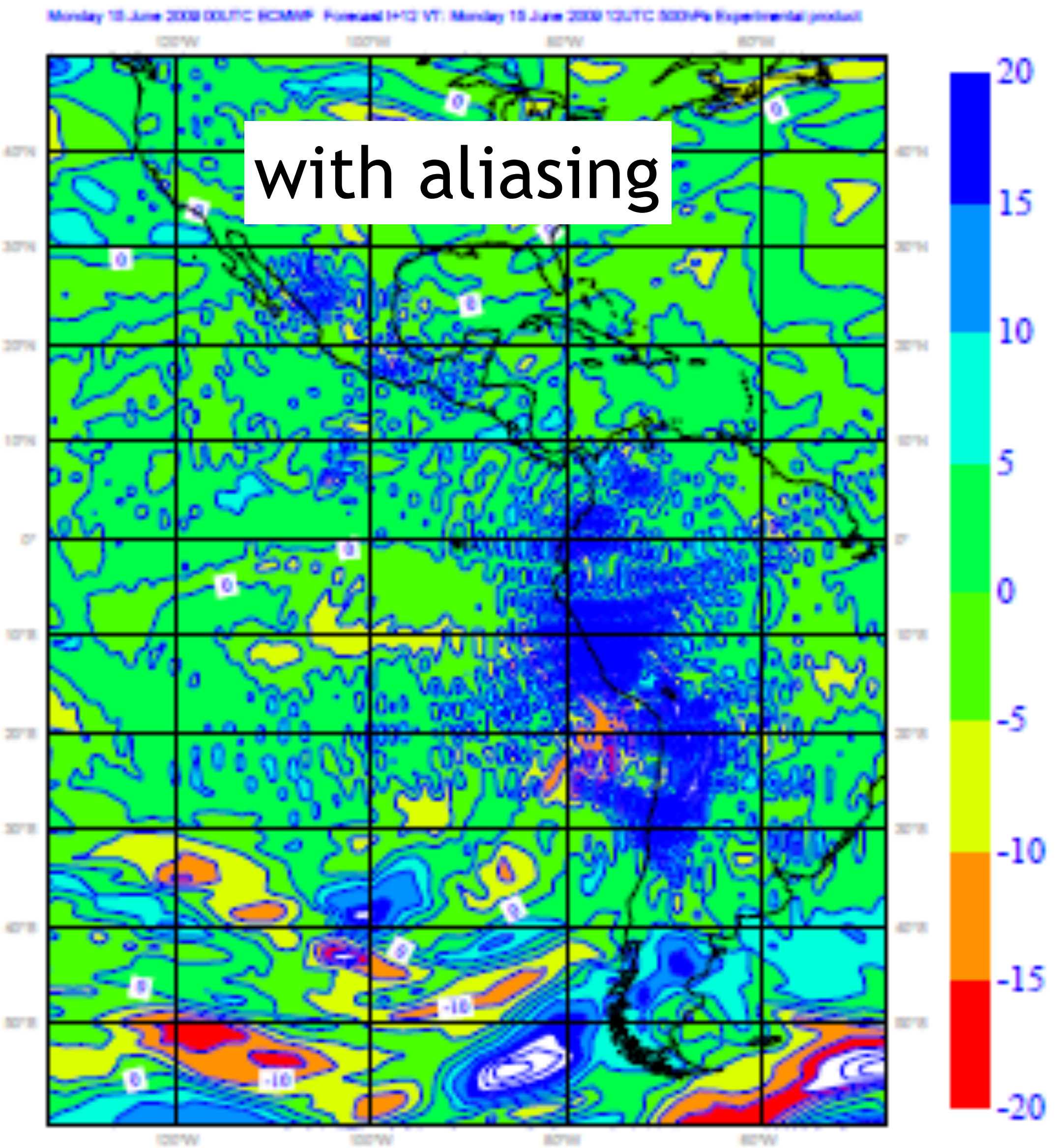
wave in grid point space

**Issue:** multiplication of two variables produces shorter waves than grid can handle



# aliasing example

500hPa adiabatic zonal wind tendencies (T159)





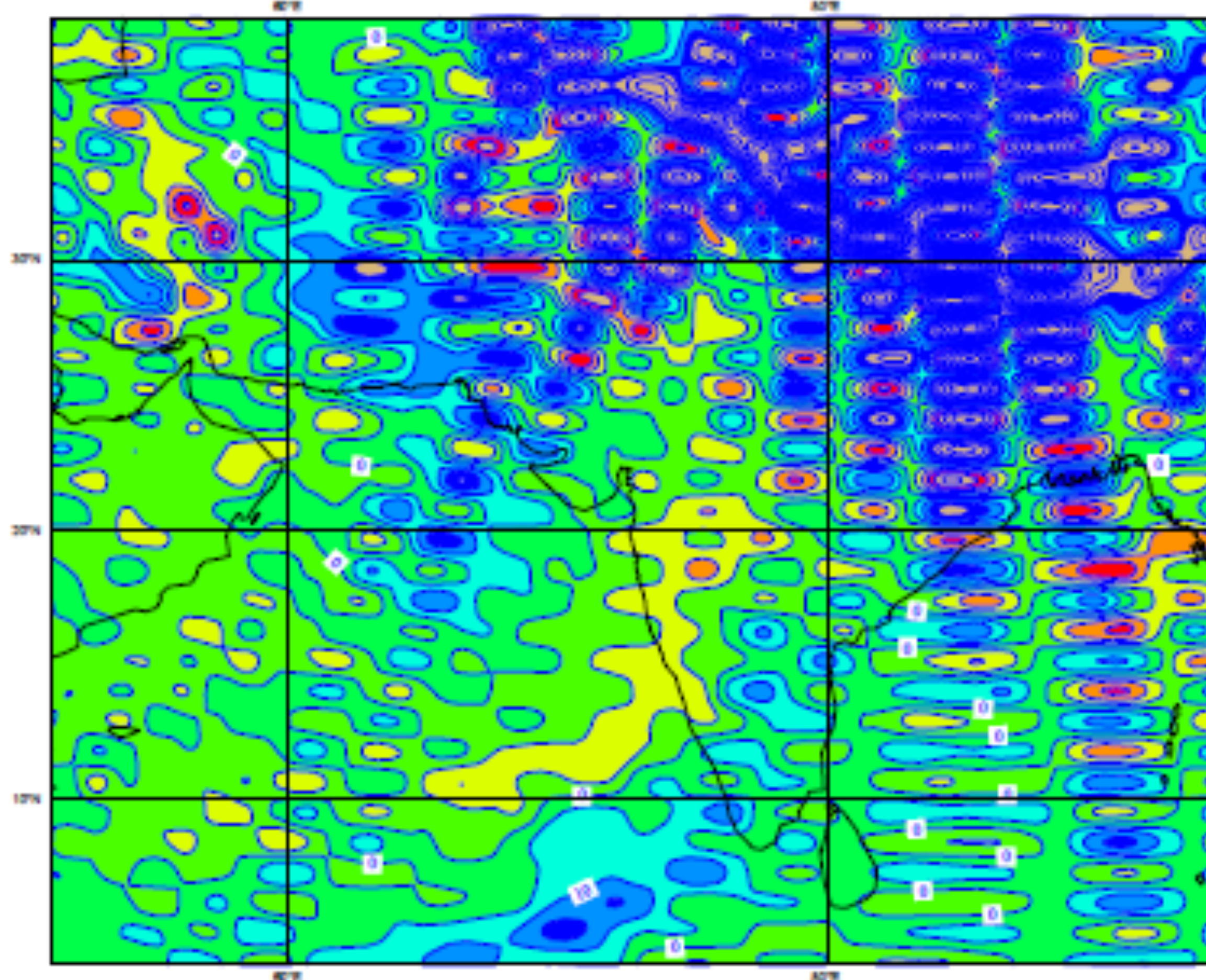
# aliasing example

500hPa adiabatic meridional wind tendencies (T159)

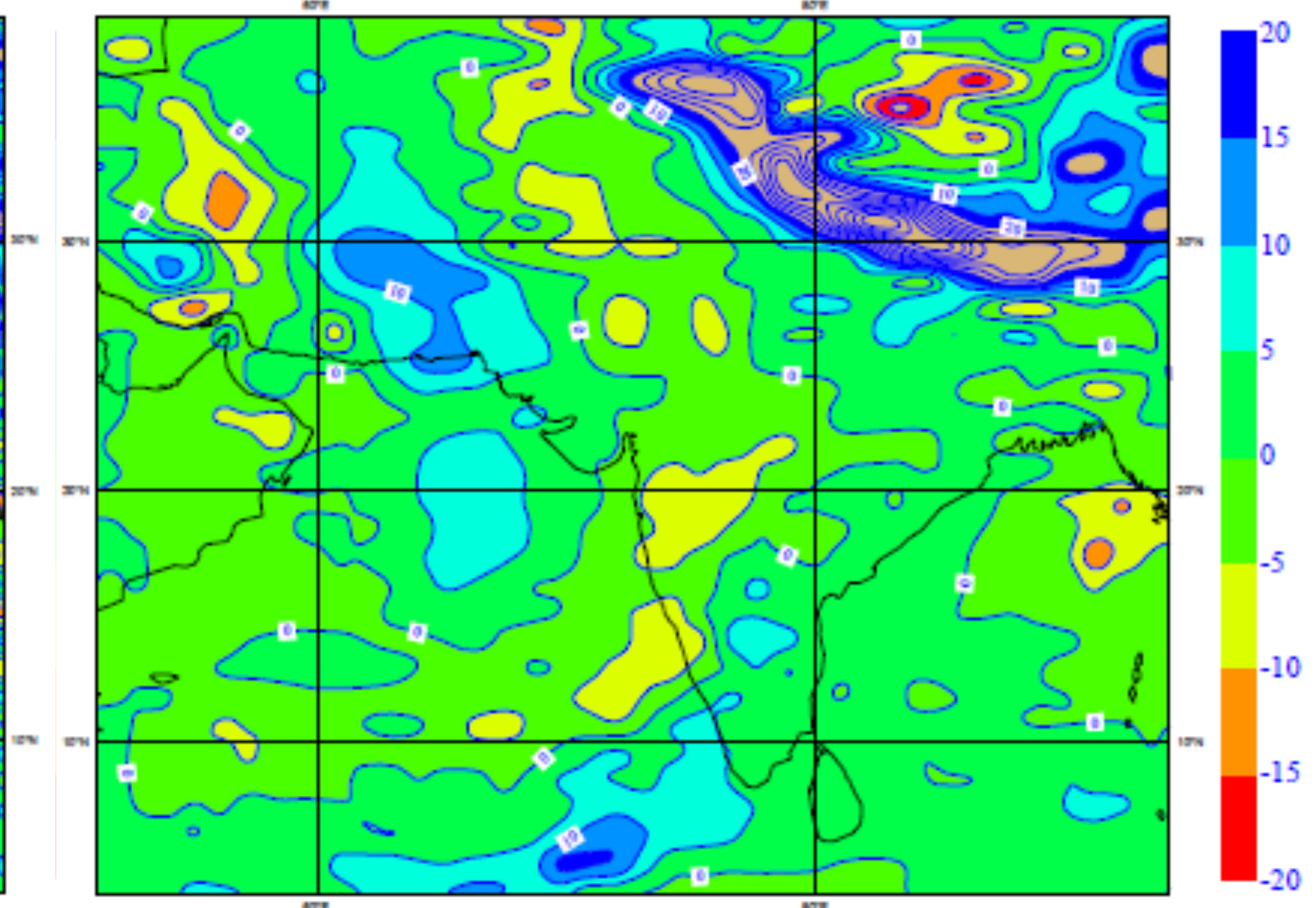
with aliasing

filtered

Monday 15 June 2009 00UTC ECMWF Forecast t+24 VT: Tuesday 16 June 2009 00UTC 500hPa Experimental product



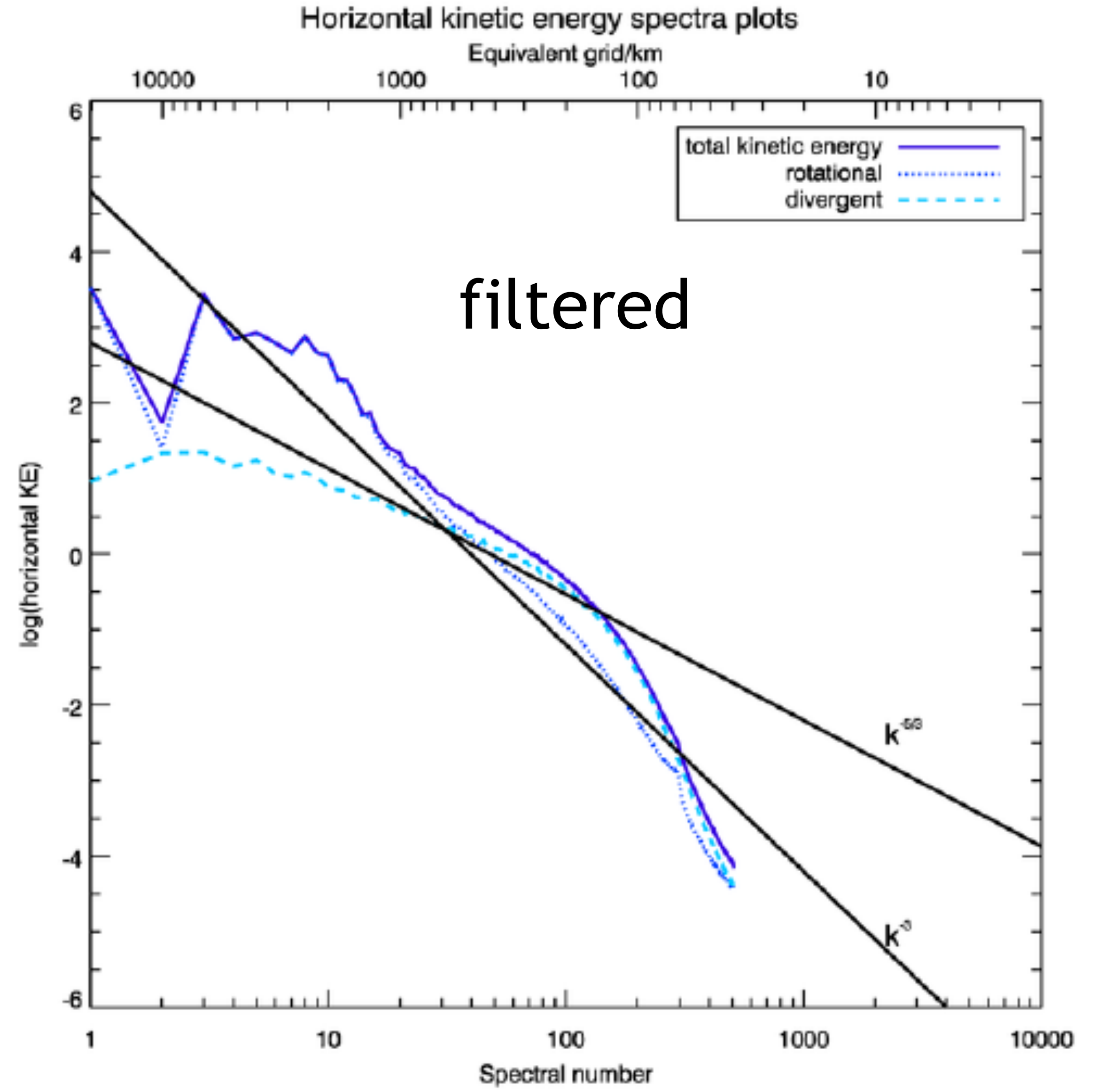
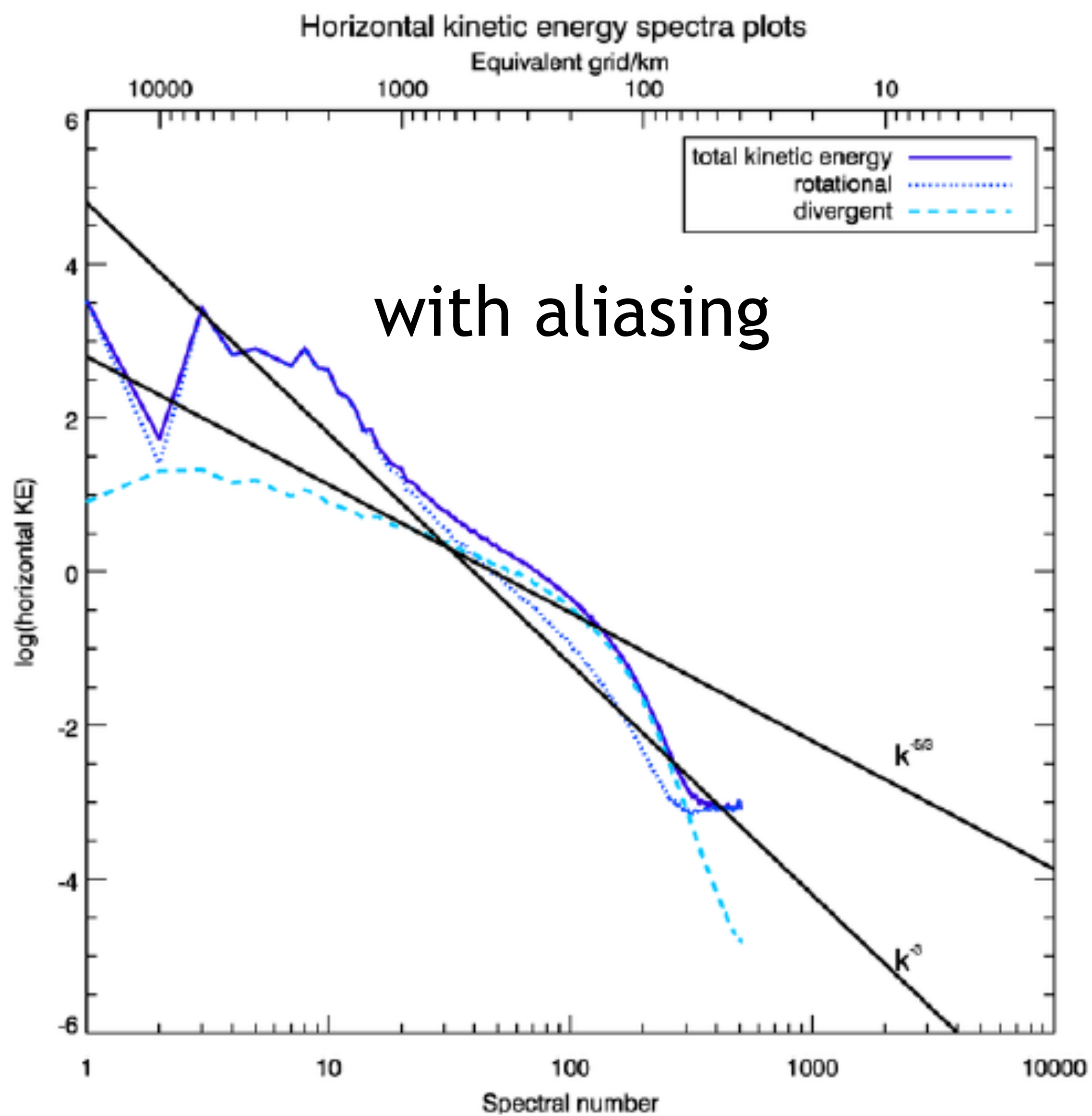
Monday 15 June 2009 00UTC ECMWF Forecast t+24 VT: Tuesday 16 June 2009 00UTC 500hPa Experimental product





# aliasing example

kinetic energy spectra, 100 hPa





# alternatives to using a filter

**Idea:** use more grid points than spectral coefficients

Orszag, 1971:

2N+1 gridpoints to N waves : linear grid ~ 1-2  $\Delta$

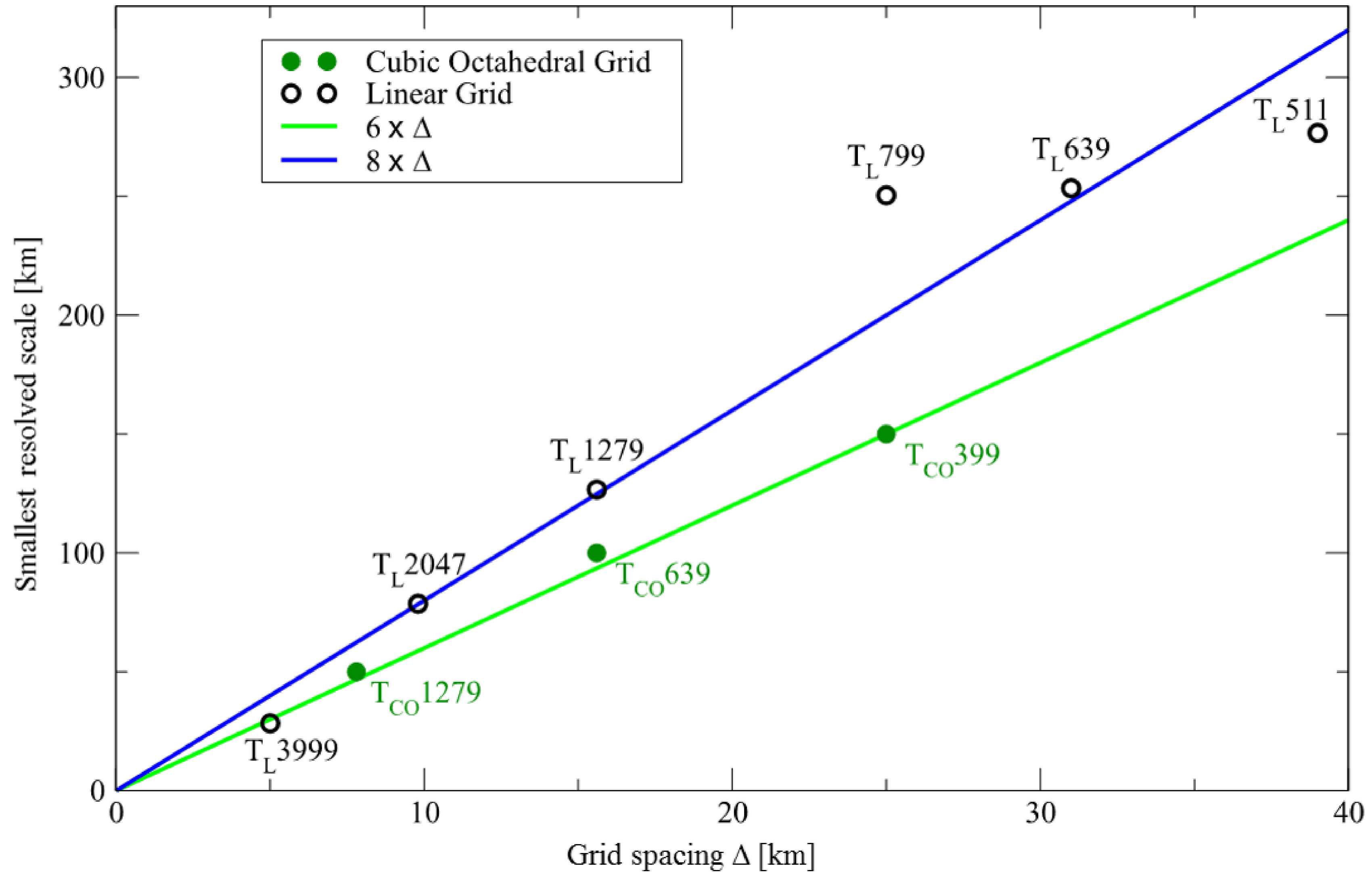
3N+1 gridpoints to N waves : quadratic grid ~ 2-3  $\Delta$

4N+1 gridpoints to N waves : cubic grid ~ 3-4  $\Delta$  (*Wedi, 2014*)

Spatial filter range



# effective resolution of linear and cubic grids (Abdalla et al. 2013)







# inverse spectral transform

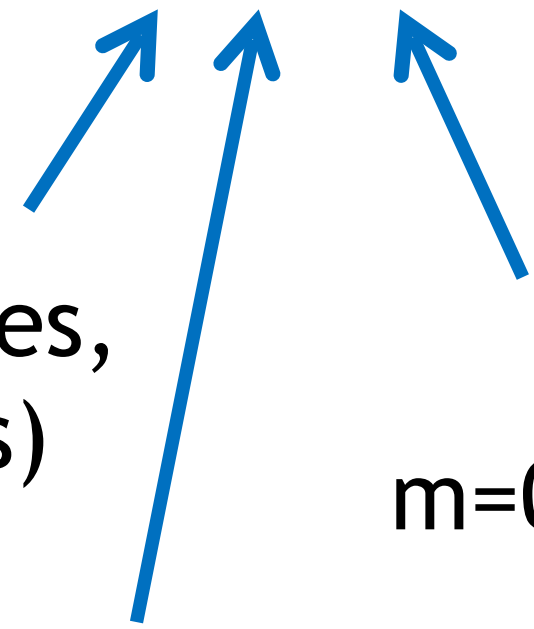
spectral data:  $\mathbf{D}(f, i, n, m)$

fastest index left (column-major  
order like in Fortran)

fields (variables,  
height levels)

wave numbers  
 $m=0, \dots, N; n=0, \dots, N-m$   
( $N$ : truncation)

real and  
imaginary part



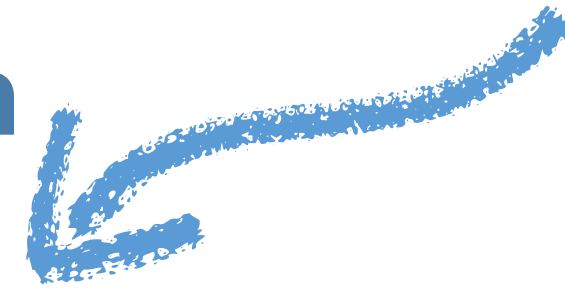


# inverse spectral transform

spectral data:  $\mathbf{D}(f, i, n, m)$

$m=0, \dots, N$ ;  $n=0, \dots, N-m$

even  $n$



odd  $n$



for each  $m$ :

$\mathbf{D}_{e,m}(f, i, n)$

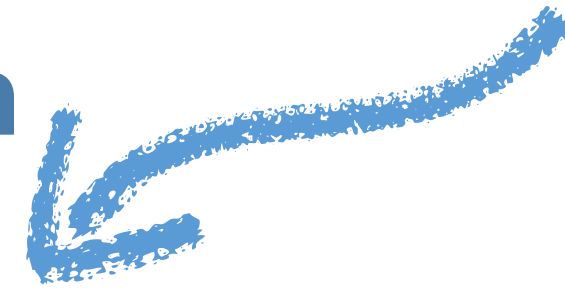
$\mathbf{D}_{o,m}(f, i, n)$



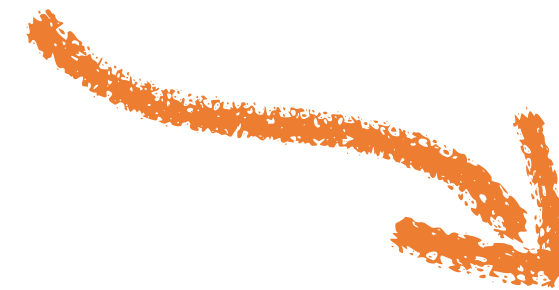
# inverse spectral transform

spectral data:  $\mathbf{D}(f, i, n, m)$

even  $n$



odd  $n$



$m=0, \dots, N; n=0, \dots, N-m$

$\mathbf{P}$ : precomputed Legendre polynomials

for each  $m$ :

$$\mathbf{S}_m(f, i, \phi) = \sum_n \mathbf{D}_{e,m}(f, i, n) \cdot \mathbf{P}_{e,m}(n, \phi), \quad \mathbf{A}_m(f, i, \phi) = \sum_n \mathbf{D}_{o,m}(f, i, n) \cdot \mathbf{P}_{o,m}(n, \phi)$$

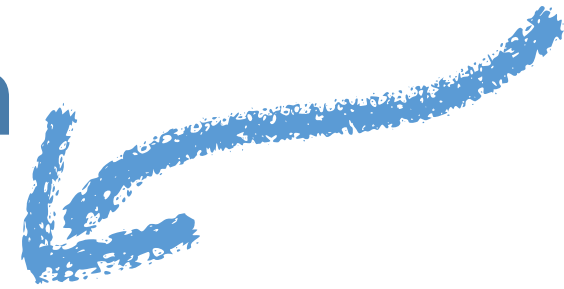
matrix  
multiplications



# inverse spectral transform

spectral data:  $\mathbf{D}(f, i, n, m)$

even  $n$



odd  $n$



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$$\phi > 0 : \mathbf{F}(i, m, \phi, f) = \mathbf{S}_m(f, i, \phi) + \mathbf{A}_m(f, i, \phi)$$

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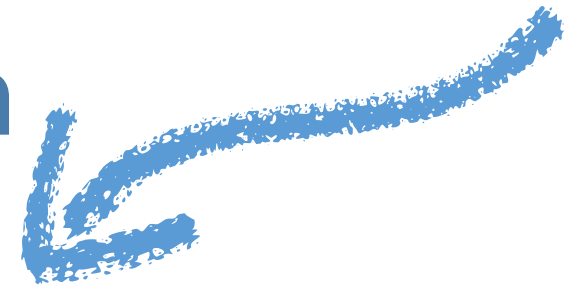
matrix  
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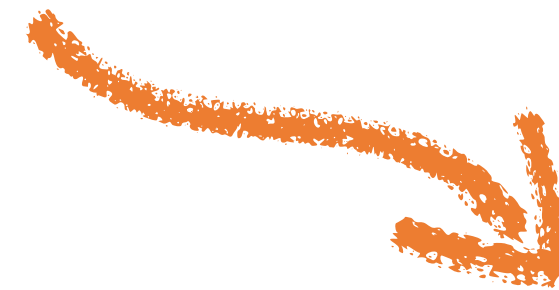
# inverse spectral transform

spectral data:  $\mathbf{D}(f, i, n, m)$

even  $n$



odd  $n$

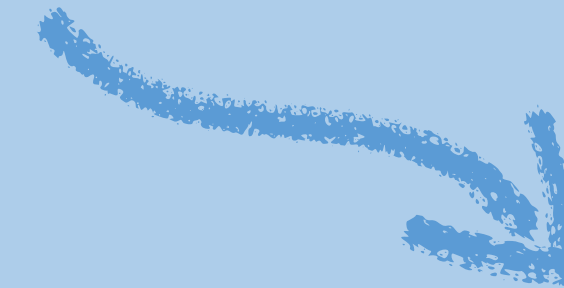


$m=0, \dots, N$ ;  $n=0, \dots, N-m$

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matrix multiplications

$$\phi > 0 : \mathbf{F}(i, m, \phi, f) = \mathbf{S}_m(f, i, \phi) + \mathbf{A}_m(f, i, \phi)$$

$$\phi < 0 : \mathbf{F}(i, m, \phi, f) = \mathbf{S}_m(f, i, -\phi) - \mathbf{A}_m(f, i, -\phi)$$

for each  $\phi, f$ :

$$\mathbf{G}_{\phi, f}(\lambda) = \text{FFT}(\mathbf{F}_{\phi, f}(i, m))$$

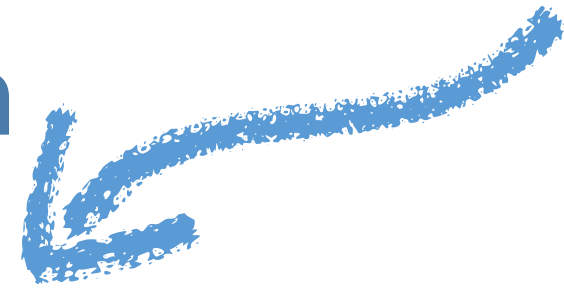
FFT: Fast Fourier Transform



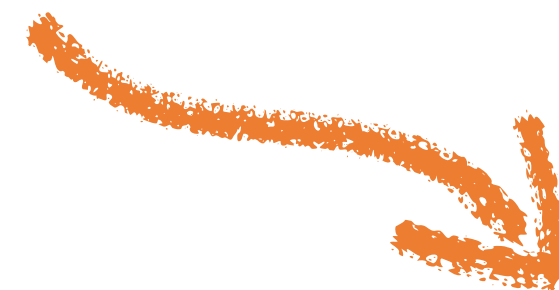
# inverse spectral transform

spectral data:  $\mathbf{D}(f, i, n, m)$

even  $n$



odd  $n$

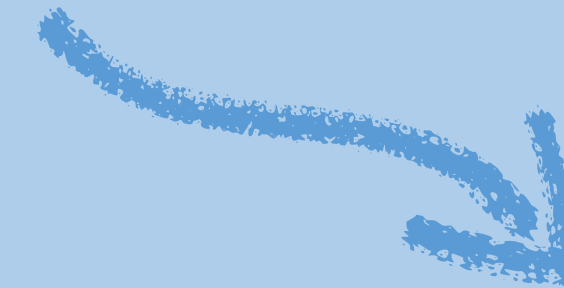


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matrix multiplications

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FFT: Fast Fourier Transform

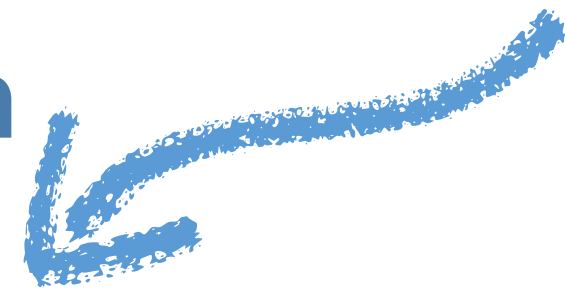
grid point data:  $\mathbf{G}(f, \lambda, \phi)$



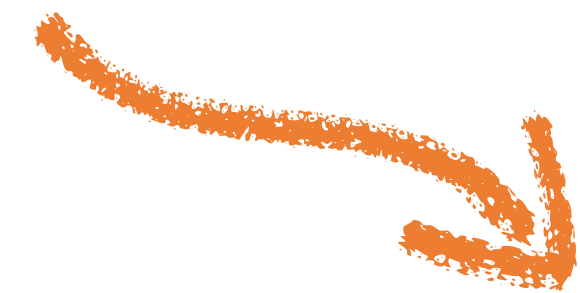
# inverse spectral transform

spectral data:  $\mathbf{D}(f, i, n, m)$

even  $n$



odd  $n$



for each  $m$ :

$$\mathbf{S}_m(f, i, \phi) = \sum_n \mathbf{D}_{e,m}(f, i, n) \cdot \mathbf{P}_{e,m}(n, \phi),$$

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spectral space

inverse Legendre transform

for each  $\phi, f$ :  $\mathbf{G}_{\phi, f}(\lambda) = \text{FFT}(\mathbf{F}_{\phi, f}(i, m))$

inverse Fourier transform

grid point data:  $\mathbf{G}(f, \lambda, \phi)$

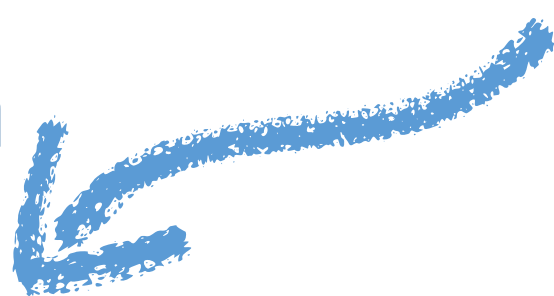
grid point space



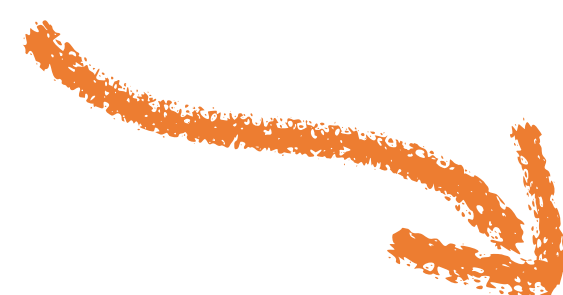
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spectral data:  $\mathbf{D}(f, i, n, m)$

even  $n$



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for each  $m$ :

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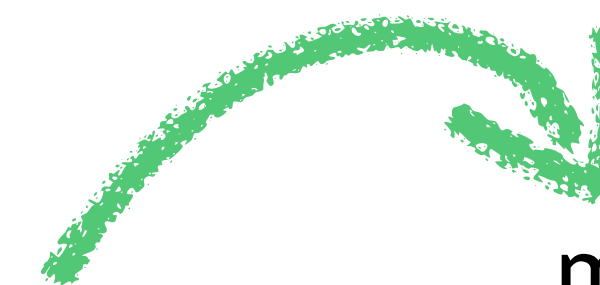
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$$\phi < 0 : \mathbf{F}(i, m, \phi, f) = \mathbf{S}_m(f, i, -\phi) - \mathbf{A}_m(f, i, -\phi)$$

for each  $\phi, f$ :  $\mathbf{G}_{\phi, f}(\lambda) = \text{FFT}(\mathbf{F}_{\phi, f}(i, m))$

grid point data:  $\mathbf{G}(f, \lambda, \phi)$

spectral space



$m, n$

parallelisation  
over these  
indices

inverse Legendre transform

$m, f$

inverse Fourier transform

$\phi, f$

grid point space

$\phi, \lambda$

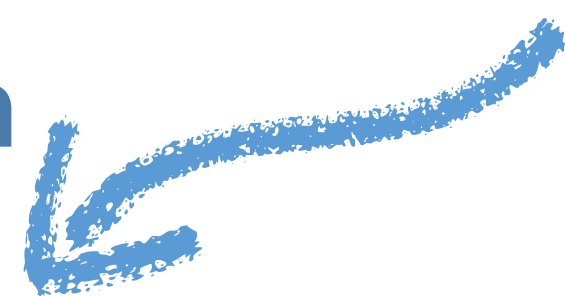




# inverse spectral transform

spectral data:  $\mathbf{D}(f, i, n, m)$

even  $n$



odd  $n$



for each  $m$ :

$$\mathbf{S}_m(f, i, \phi) = \sum_n \mathbf{D}_{e,m}(f, i, n) \cdot \mathbf{P}_{e,m}(n, \phi),$$

$$\mathbf{A}_m(f, i, \phi) = \sum_n \mathbf{D}_{o,m}(f, i, n) \cdot \mathbf{P}_{o,m}(n, \phi)$$

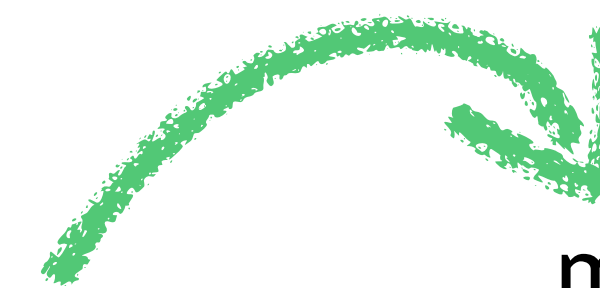
$$\phi > 0 : \mathbf{F}(i, m, \phi, f) = \mathbf{S}_m(f, i, \phi) + \mathbf{A}_m(f, i, \phi)$$

$$\phi < 0 : \mathbf{F}(i, m, \phi, f) = \mathbf{S}_m(f, i, -\phi) - \mathbf{A}_m(f, i, -\phi)$$

for each  $\phi, f$ :  $\mathbf{G}_{\phi, f}(\lambda) = \text{FFT}(\mathbf{F}_{\phi, f}(i, m))$

grid point data:  $\mathbf{G}(f, \lambda, \phi)$

spectral space



$m, n$

parallelisation  
over these  
indices

lots of MPI  
communication

inverse Legendre transform

$m, f$

inverse Fourier transform

$\phi, f$

grid point space

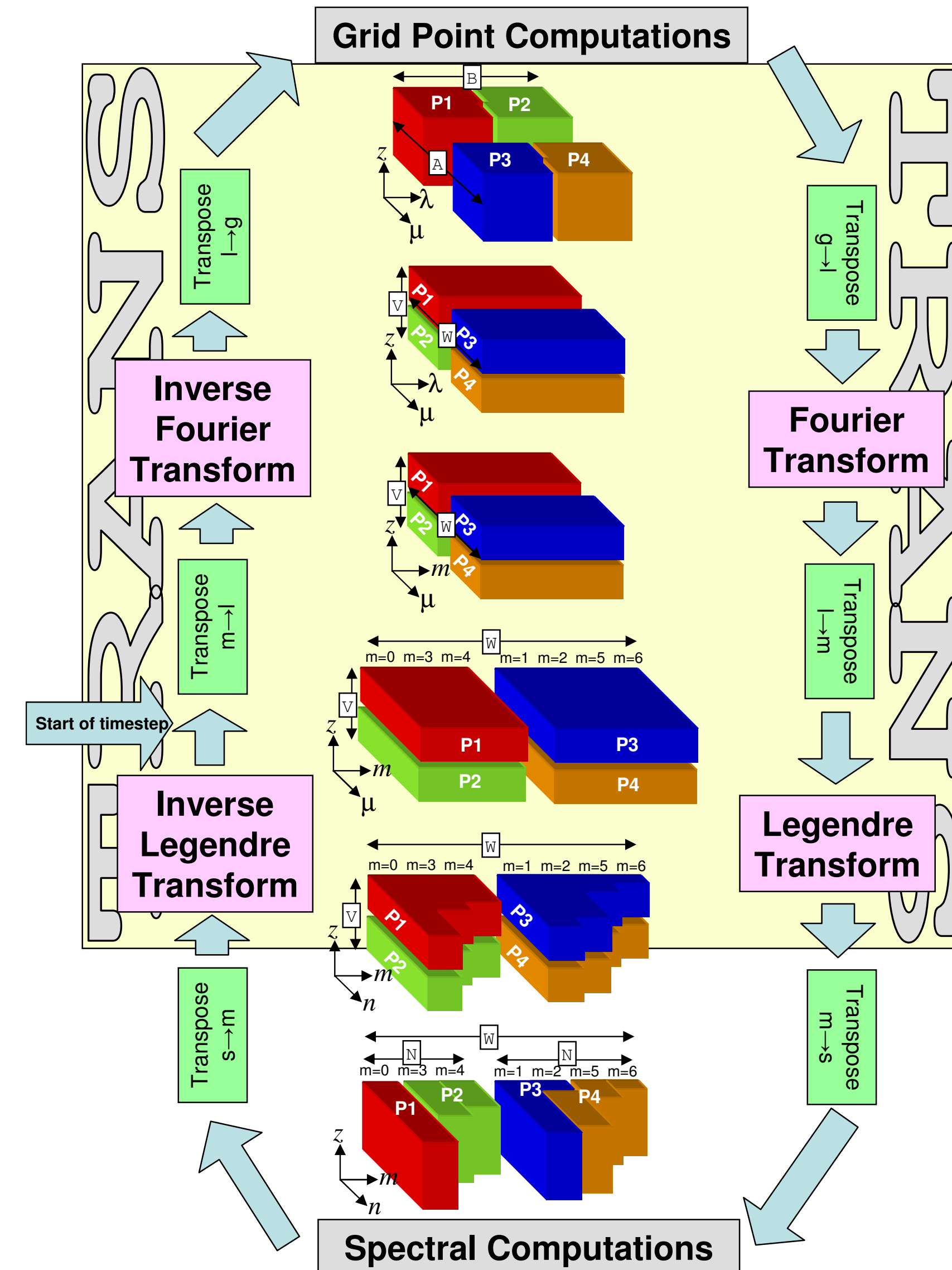
$\phi, \lambda$





# direct spectral transform

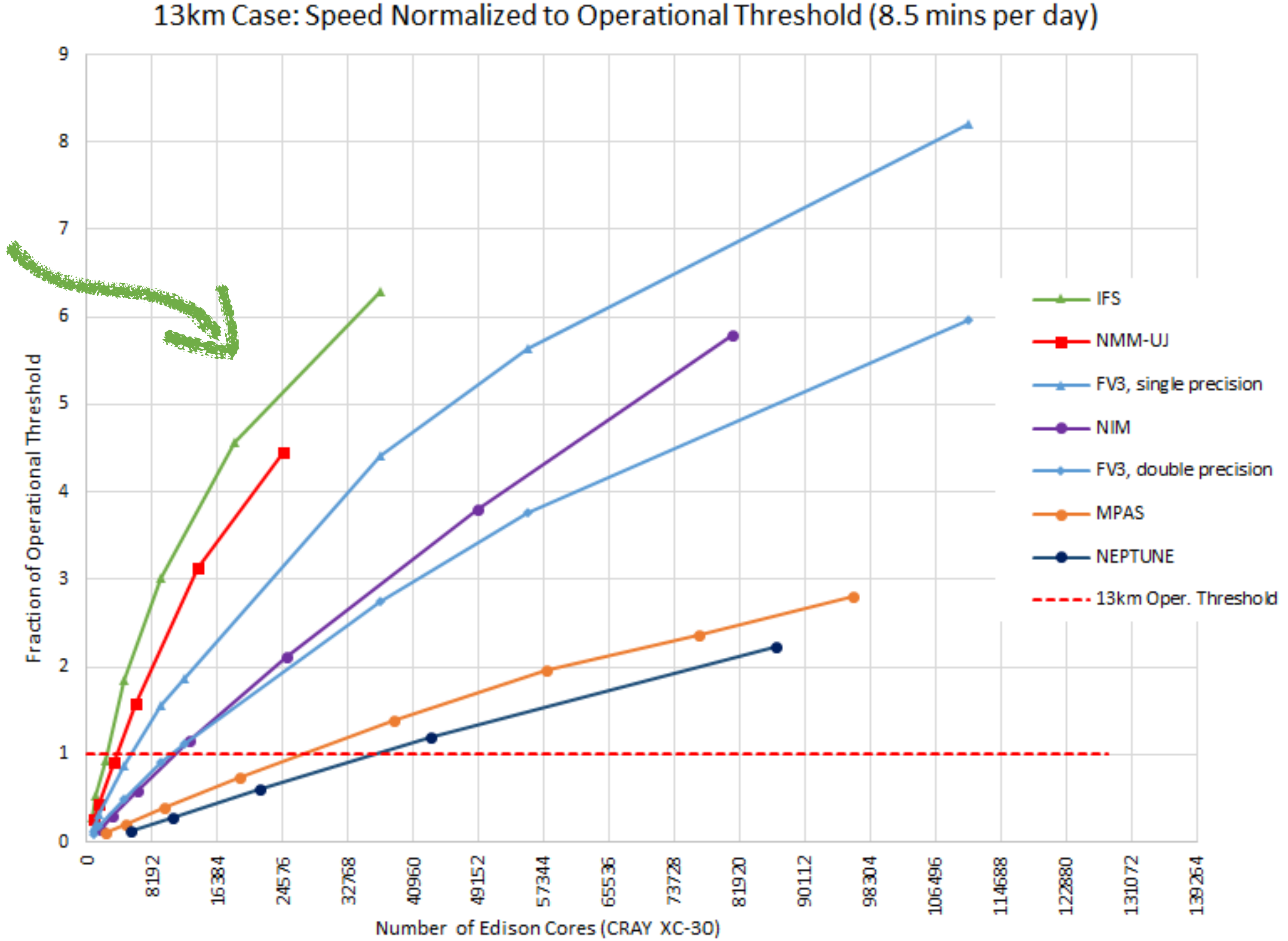
- same like inverse spectral transform
- reverse order
- multiply data with Gaussian quadrature weights before Legendre transform





# performance comparison of IFS with other models

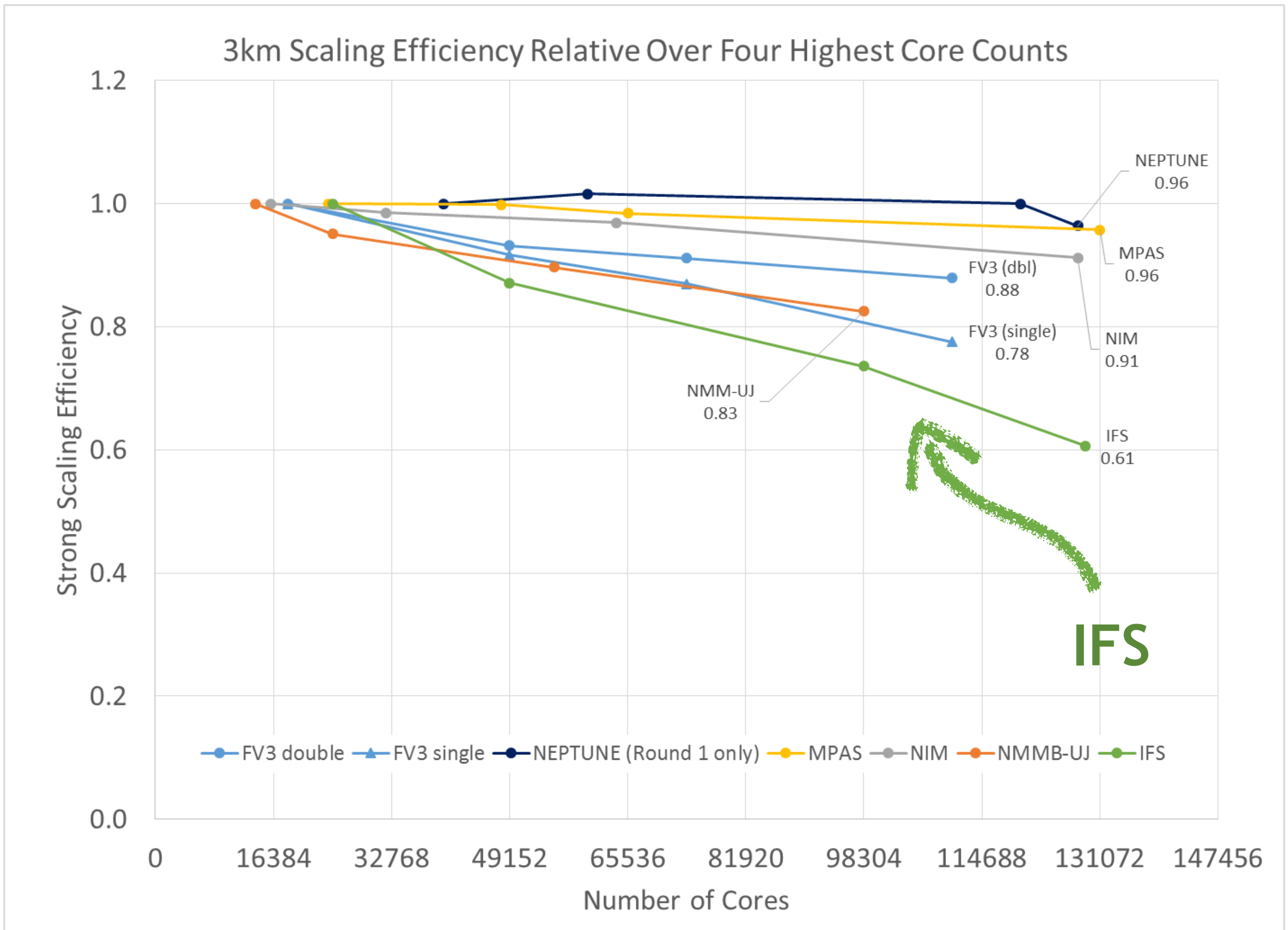
IFS



(Michalakes et al, NGGPS AVEC report, 2015)



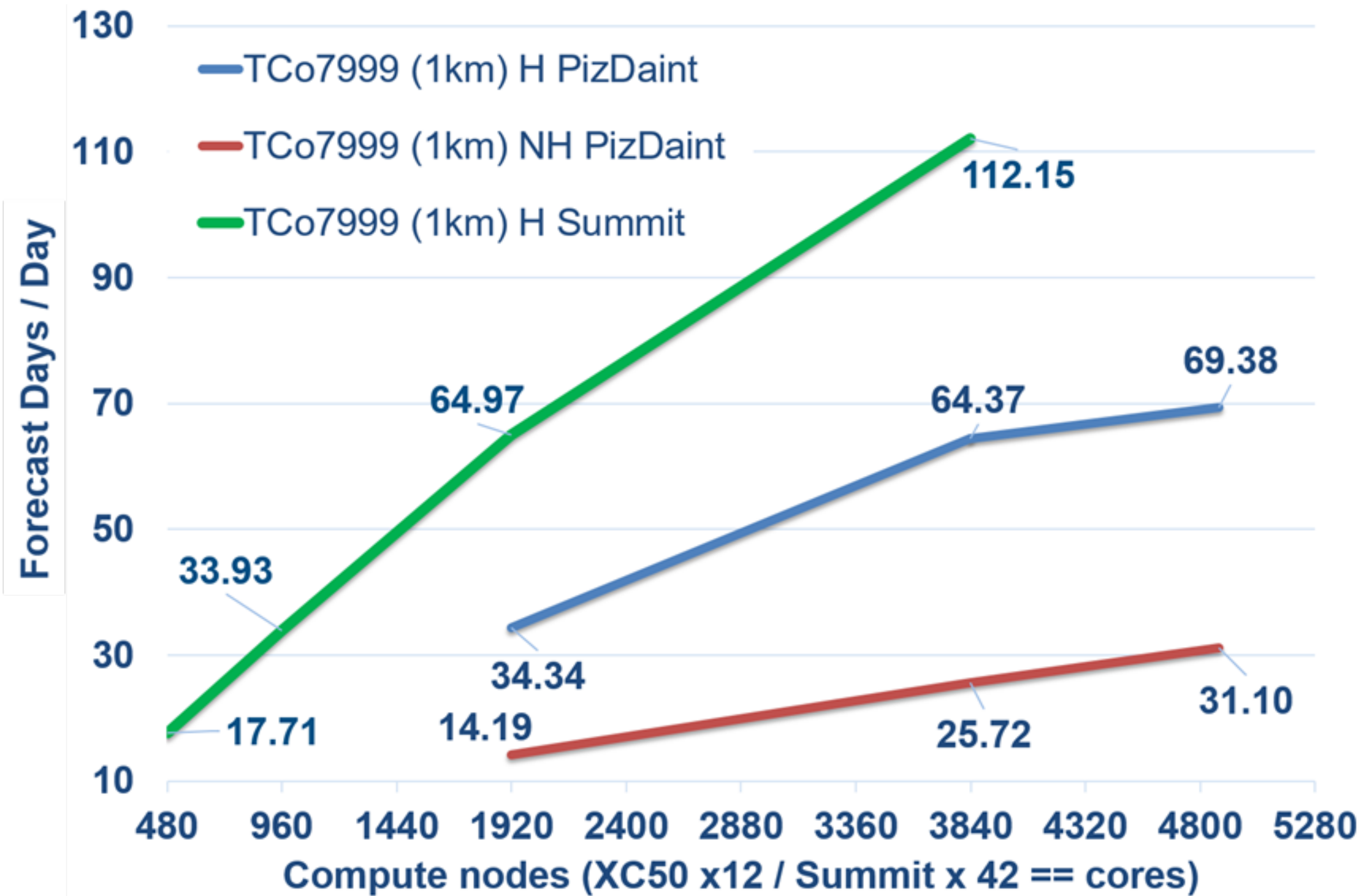
# scalability comparison of IFS with other models



IFS

(Michalakes et al, NGGPS AVEC report, 2015)

# IFS scaling on Summit and PizDaint (CPU only)



# spectral transform vs discontinuous Galerkin

projected for 5km 2-day forecast



DG, horizontally  
explicit => 4s time-  
step, almost no  
communication

communication  
volume:

**34 TB on  
2880 MPI procs**

---

time to solution:

**4 hours**

# spectral transform vs discontinuous Galerkin projected for 5km 2-day forecast



DG, horizontally  
explicit => 4s time-  
step, almost no  
communication

IFS (spectral  
transform): 240s  
time-step, lots of  
communication

communication  
volume:

**34 TB on  
2880 MPI procs**

**427 TB on  
2880 MPI procs**

time to solution:

**4 hours**

# spectral transform vs discontinuous Galerkin projected for 5km 2-day forecast



DG, horizontally  
explicit => 4s time-  
step, almost no  
communication

IFS (spectral  
transform): 240s  
time-step, lots of  
communication

communication  
volume:

**34 TB on  
2880 MPI procs**

**427 TB on  
2880 MPI procs**

time to solution:

**4 hours**

**12 minutes**



# spectral transform vs discontinuous Galerkin

projected for 5km 2-day forecast



DG, horizontally explicit => 4s time-step, almost no communication

IFS (spectral transform): 240s time-step, lots of communication

DG (like on the left)

communication volume:

**34 TB on 2880 MPI procs**

**427 TB on 2880 MPI procs**

**689 TB on 57600 MPI procs**

time to solution:

**4 hours**

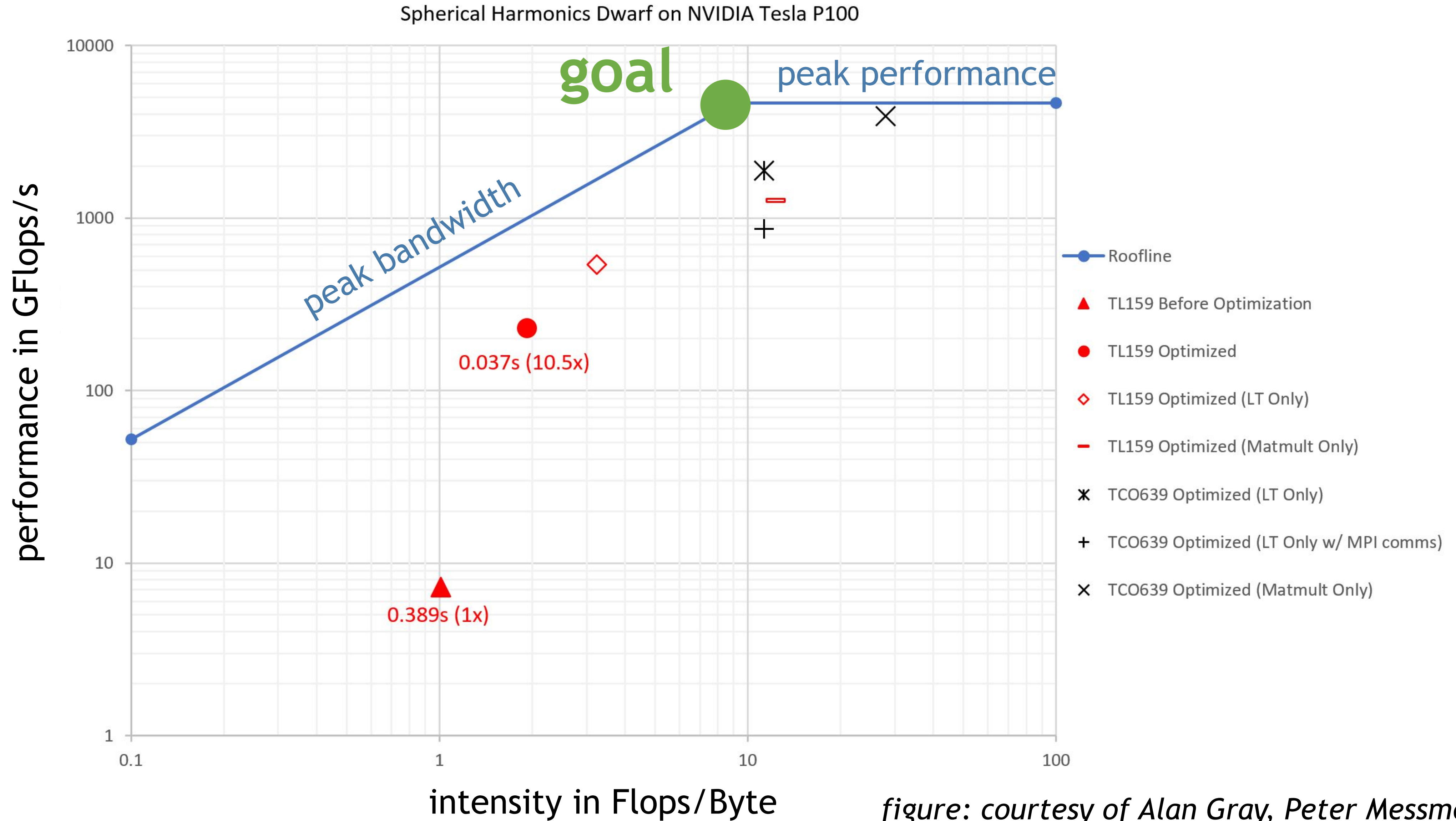
**12 minutes**

**12 minutes**



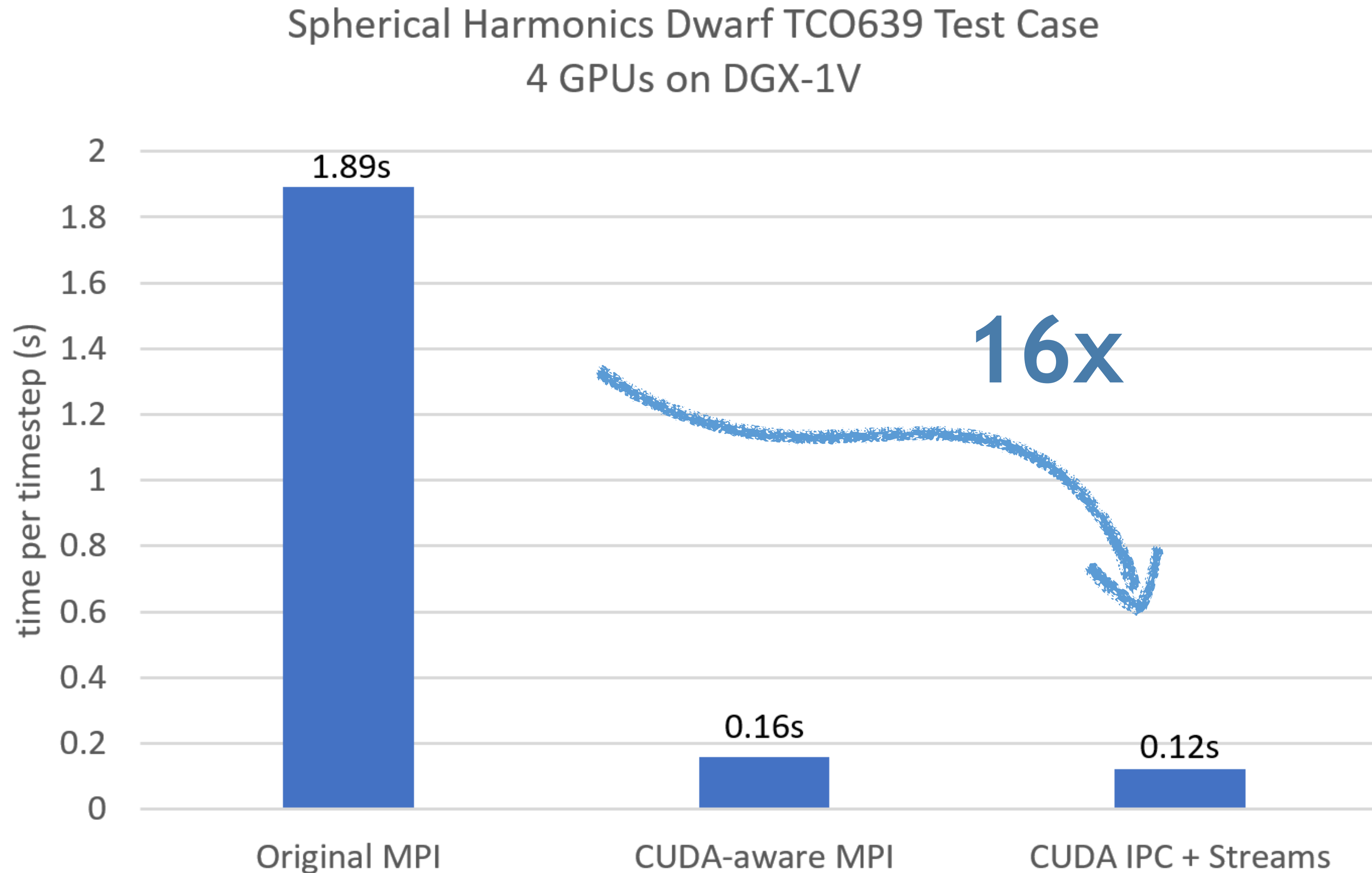


# optimisations by NVIDIA in ESCAPE





# optimisations by NVIDIA in ESCAPE

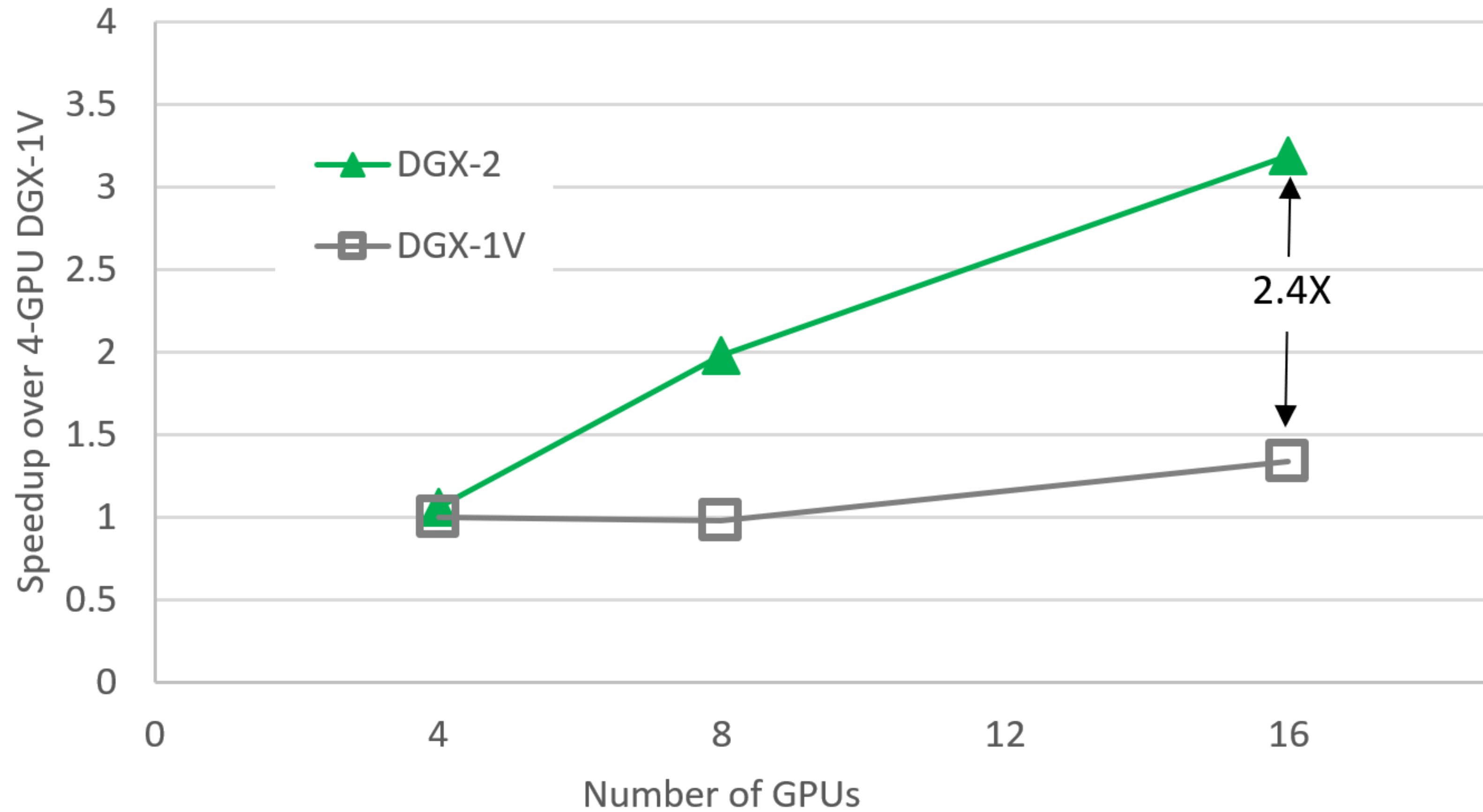


*figure: courtesy of Alan Gray, Peter Messmer (NVIDIA)*



# optimisations by NVIDIA in ESCAPE

Spherical Harmonics Dwarf TCO639 Test Case  
DGX-2 vs DGX-1V

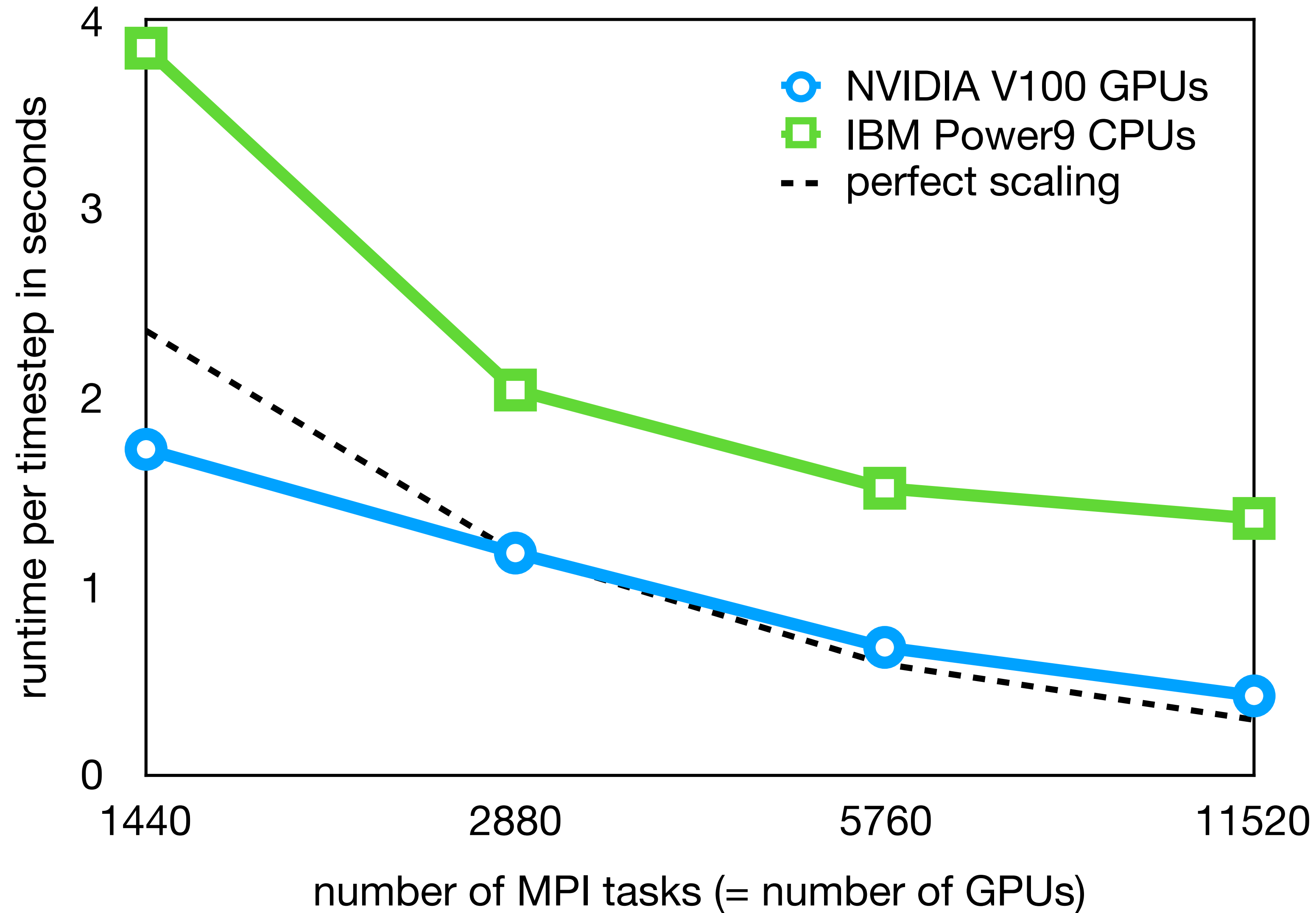


DGX-1V uses MPI for  $\geq 8$  GPUs (due to lack of AlltoAll links), all others use CUDA IPC.  
DGX-2 results use pre-production hardware.

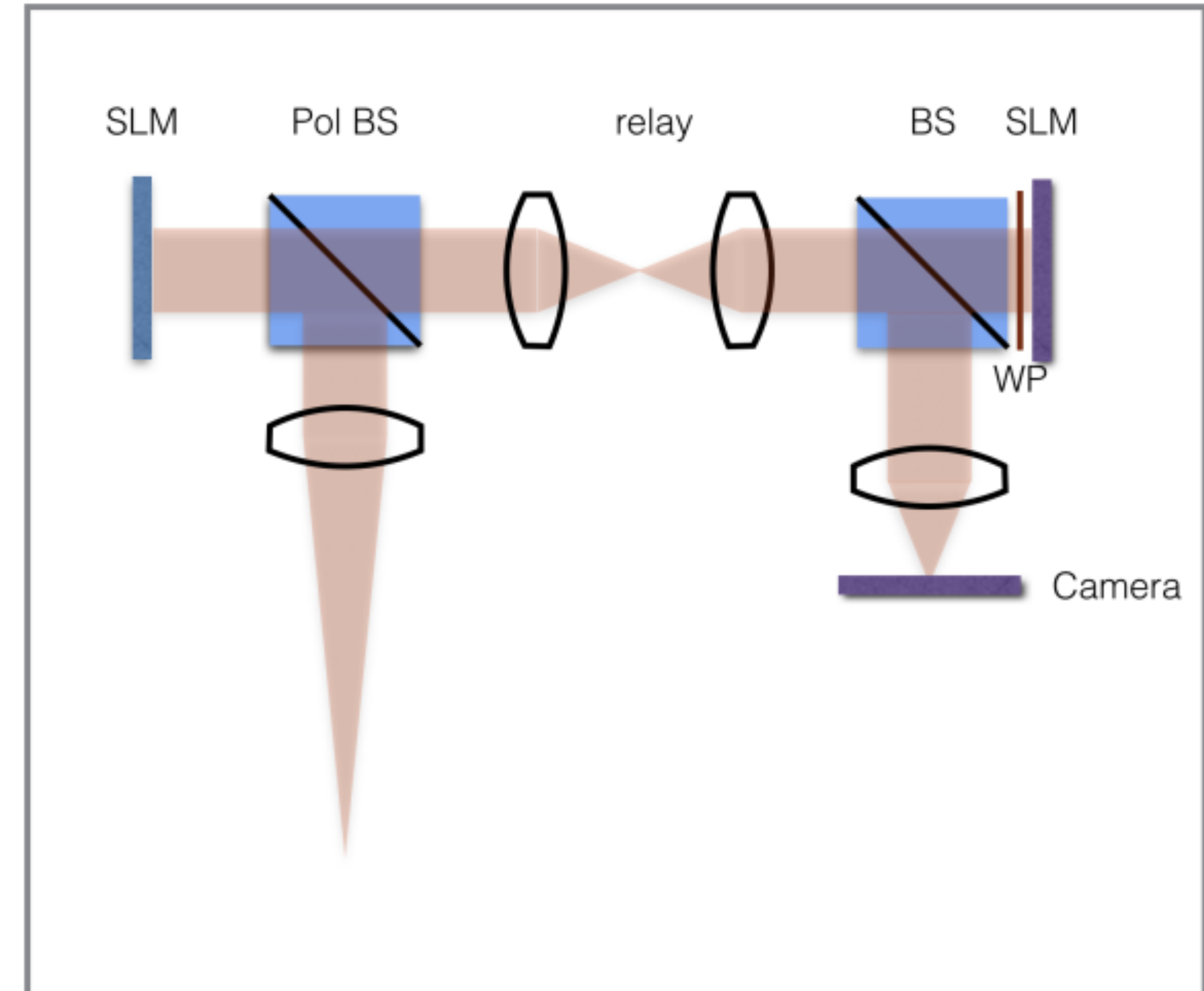
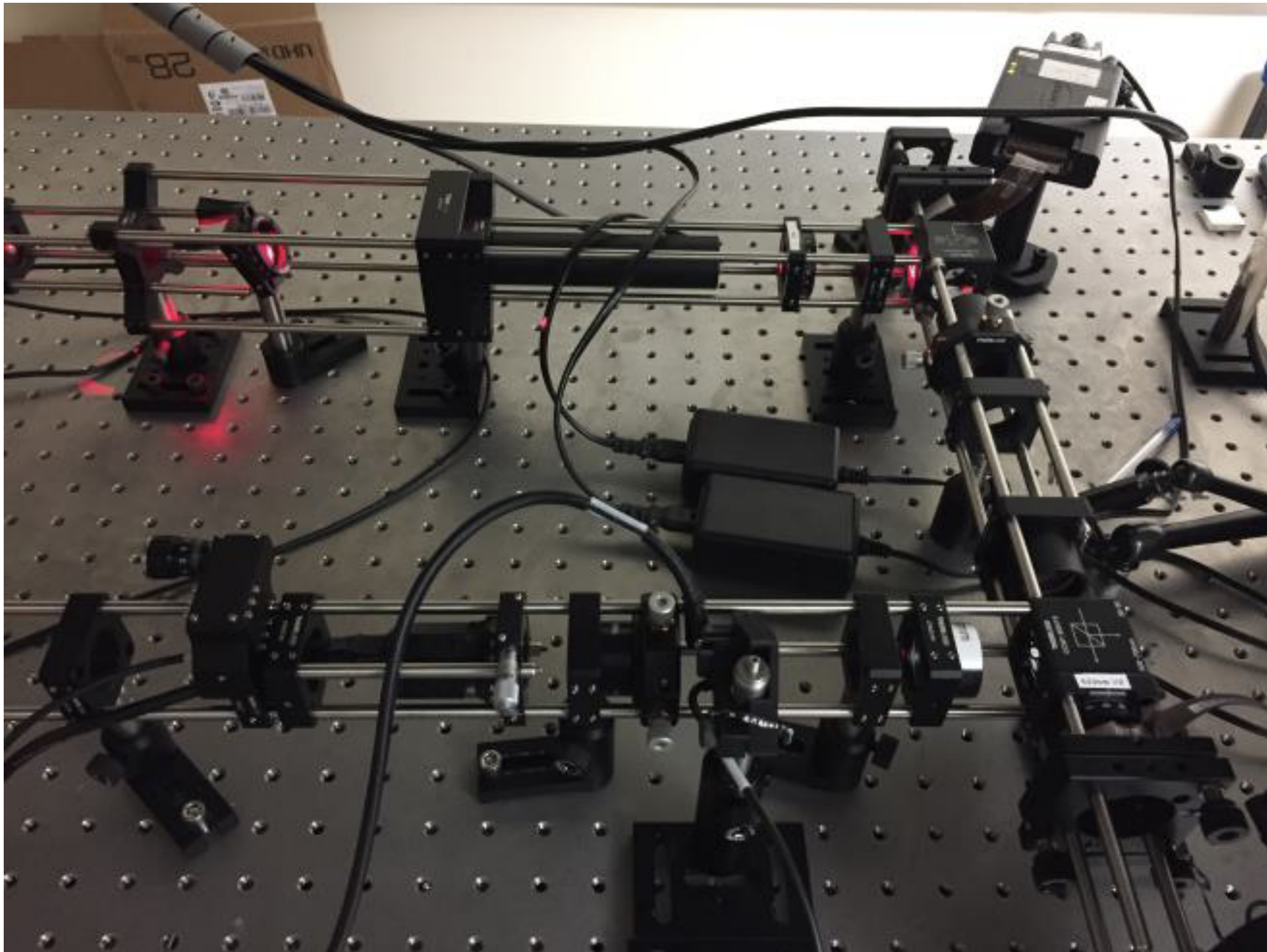
*figure: courtesy of Alan Gray,  
Peter Messmer (NVIDIA)*



# GPUs vs CPUs on Summit



# Optalysys: optical processor for spectral transform





# Fast Legendre Transform

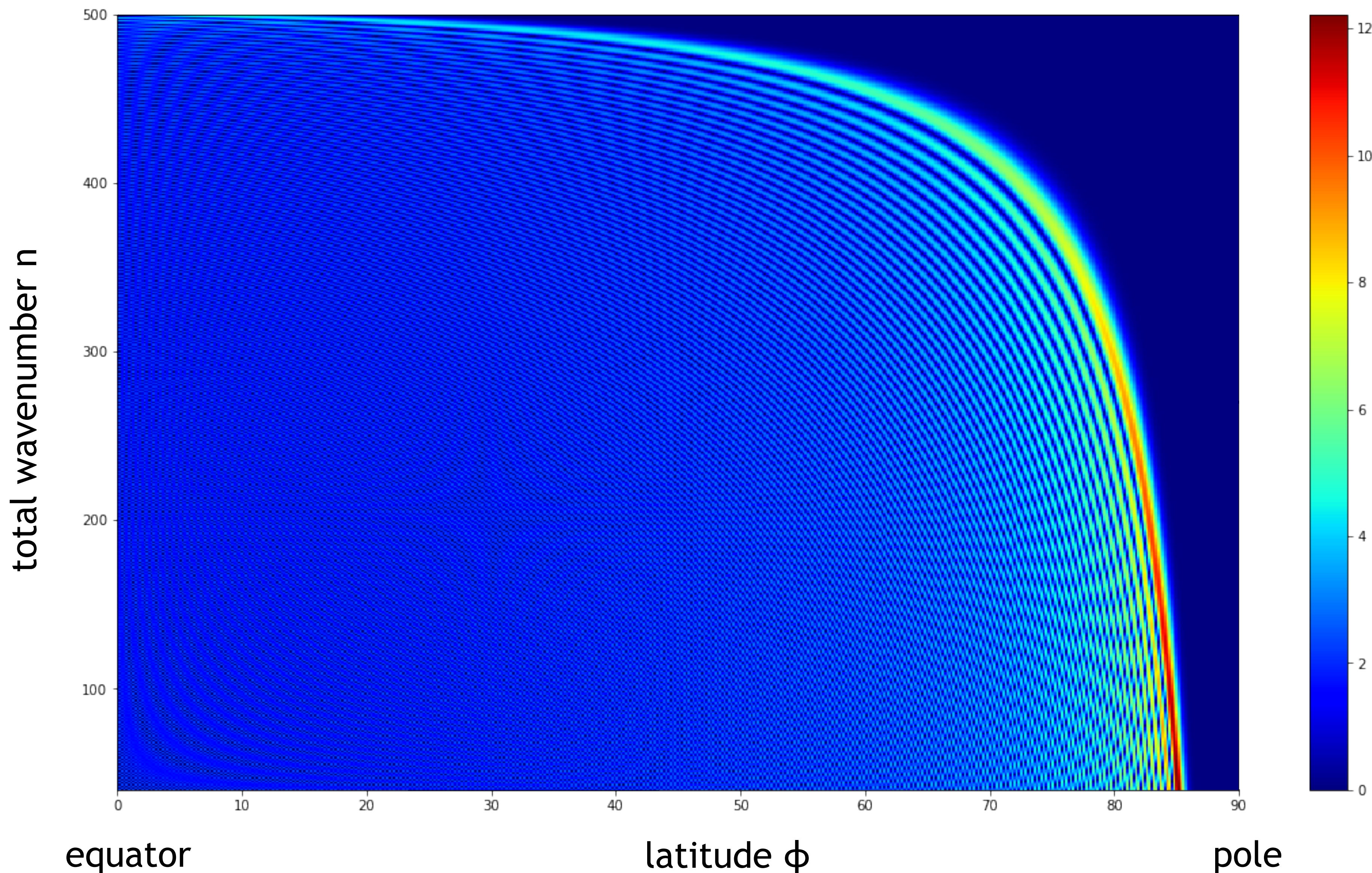
matrix of  
Legendre polynomials

truncation  $N=500$ ,  
zonal wavenumber  
 $m=40$

**FLT:**

**step 1:** split matrix  
into two rows

**step 2:** use  
interpolation to  
empty half of the  
columns





# Fast Legendre Transform

matrix of Legendre polynomials

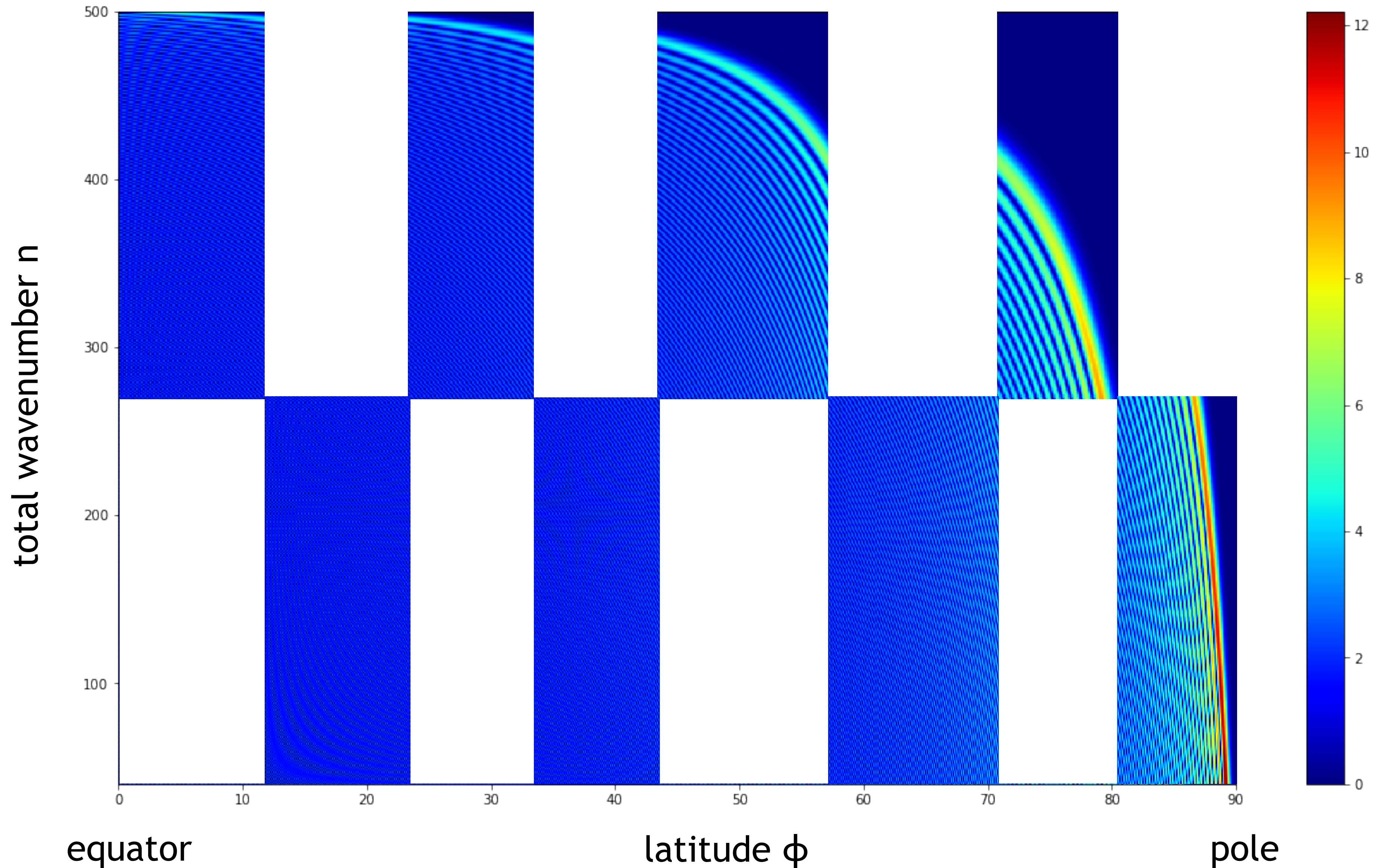
truncation  $N=500$ ,  
zonal wavenumber  $m=40$

**FLT:**

**step 1:** split matrix into two rows

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**step 3:** reorder columns







# Fast Legendre Transform

matrix of Legendre polynomials

truncation  $N=500$ ,  
zonal wavenumber  $m=40$

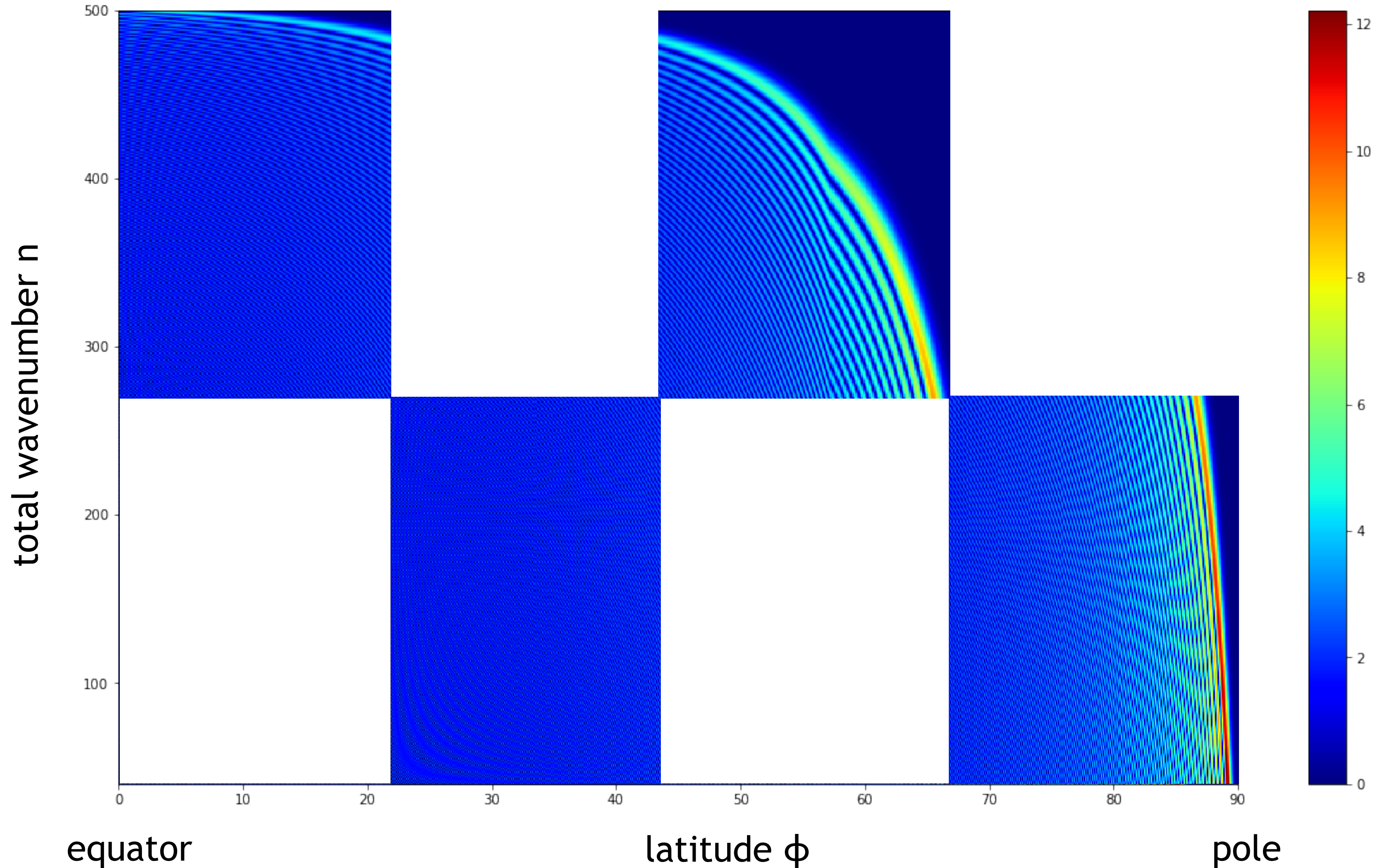
**FLT:**

**step 1:** split matrix into two rows

**step 2:** use interpolation to empty half of the columns

**step 3:** reorder columns

**step 4:** apply to each block recursively





# Fast Legendre Transform

matrix of Legendre polynomials

truncation  $N=500$ ,  
zonal wavenumber  $m=40$

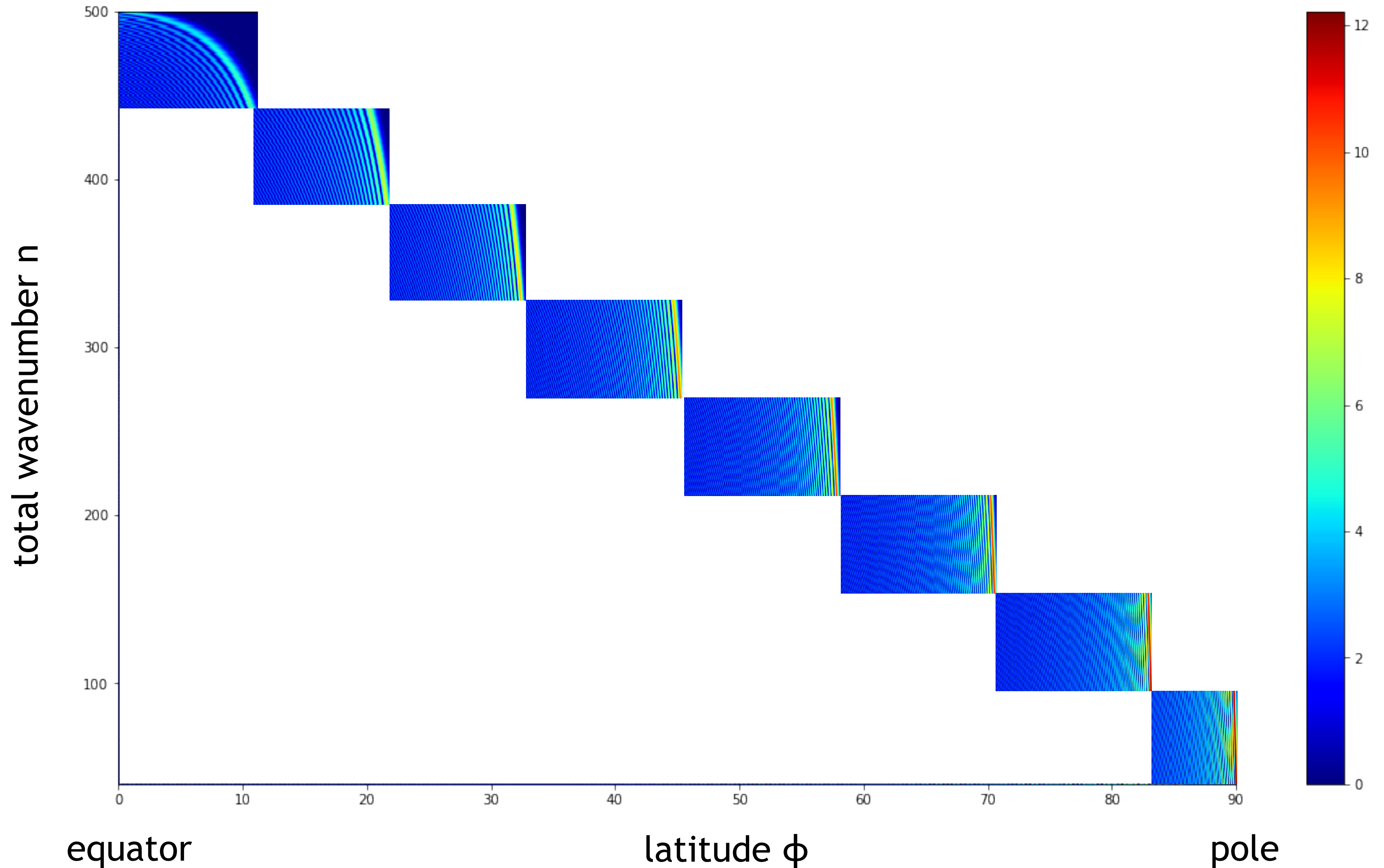
**FLT:**

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# Fast Legendre Transform

matrix of Legendre polynomials

truncation  $N=500$ ,  
zonal wavenumber  
 $m=40$

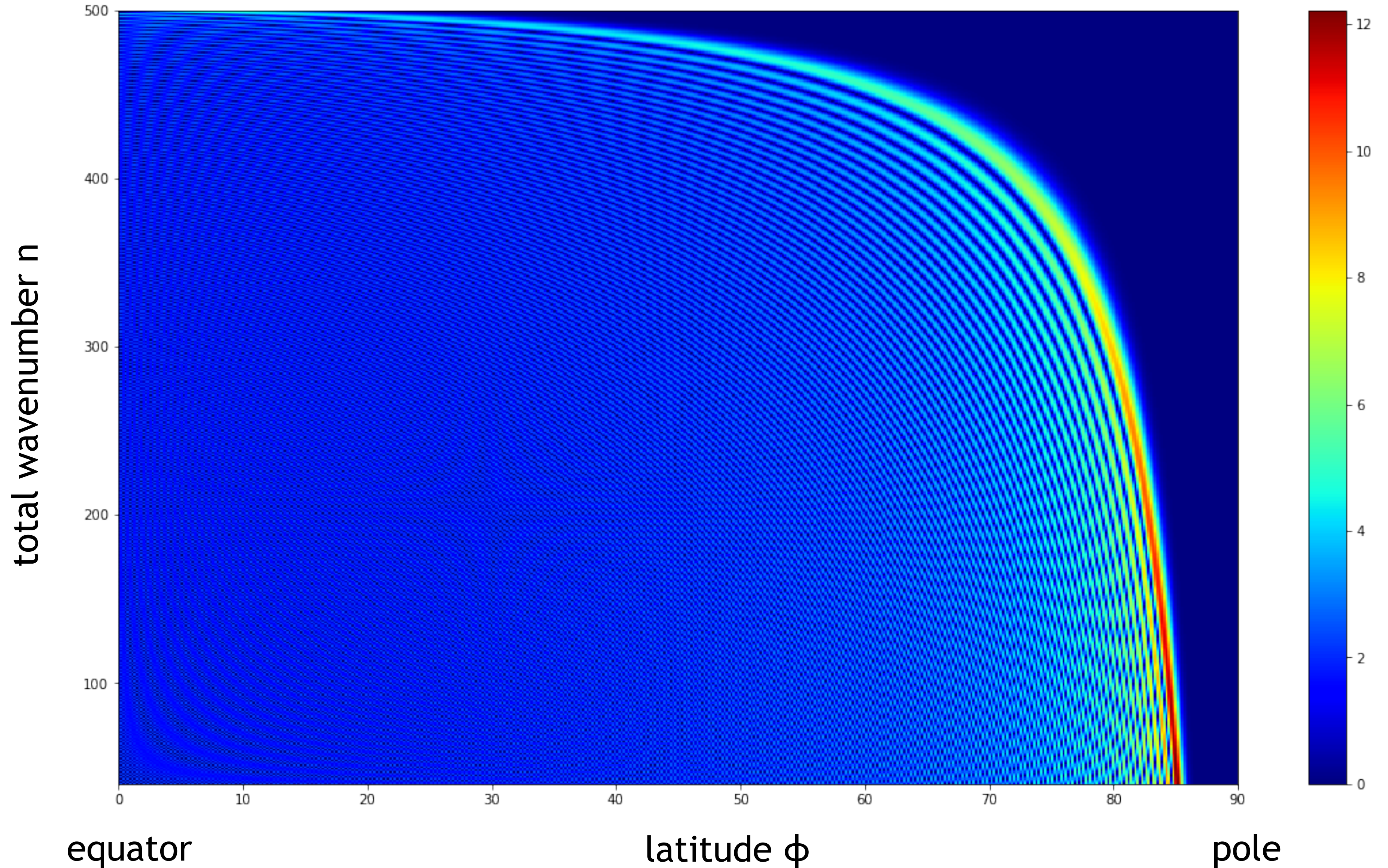
**FLT:**

**step 1:** split matrix into two rows

**step 2:** use interpolation to empty half of the columns

**step 3:** reorder columns

**step 4:** apply to each block recursively





# Fast Legendre Transform

matrix of  
Legendre polynomials

truncation  $N=500$ ,  
zonal wavenumber  
 $m=100$

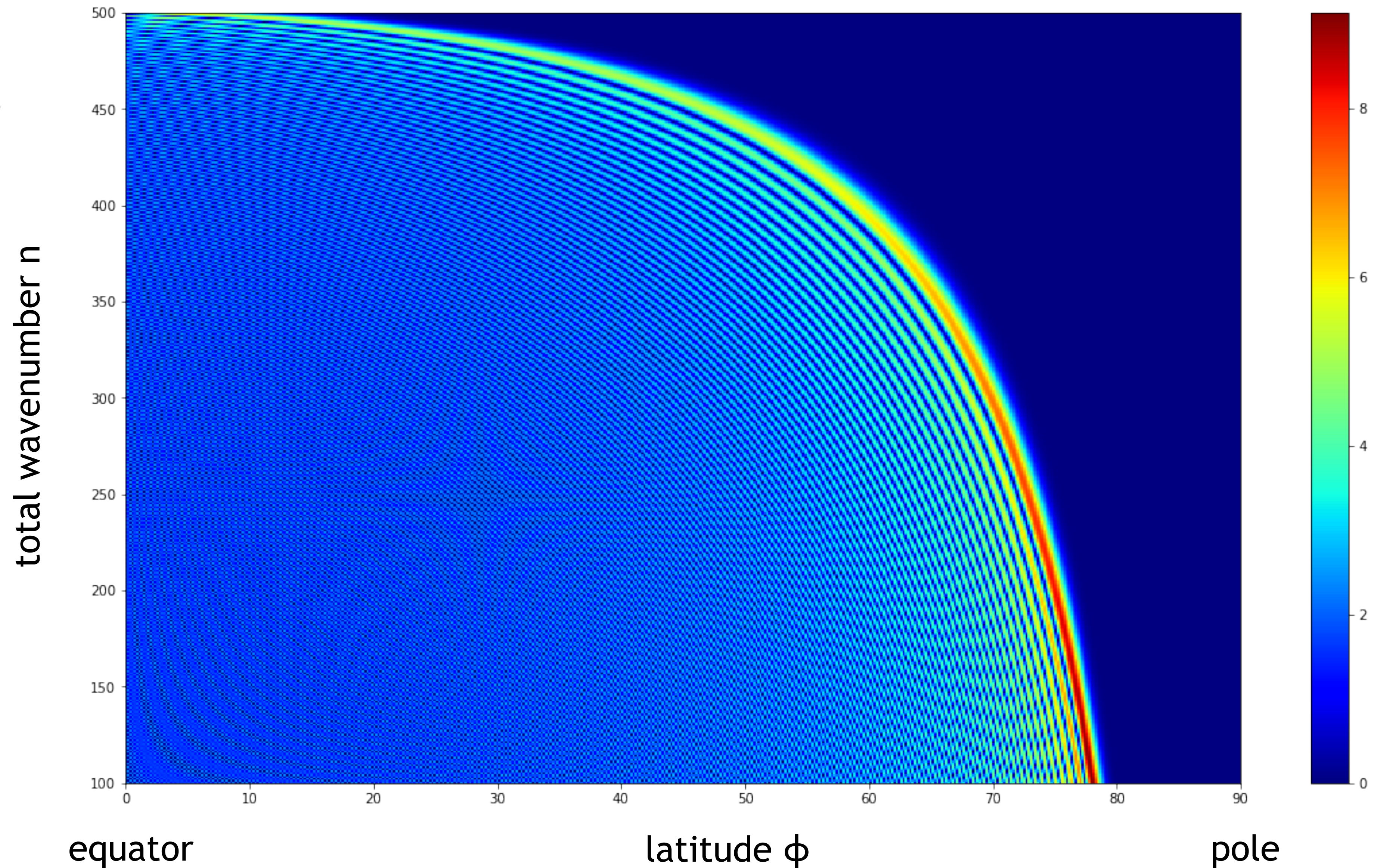
**FLT:**

**step 1:** split matrix  
into two rows

**step 2:** use  
interpolation to  
empty half of the  
columns

**step 3:** reorder  
columns

**step 4:** apply to each  
block recursively

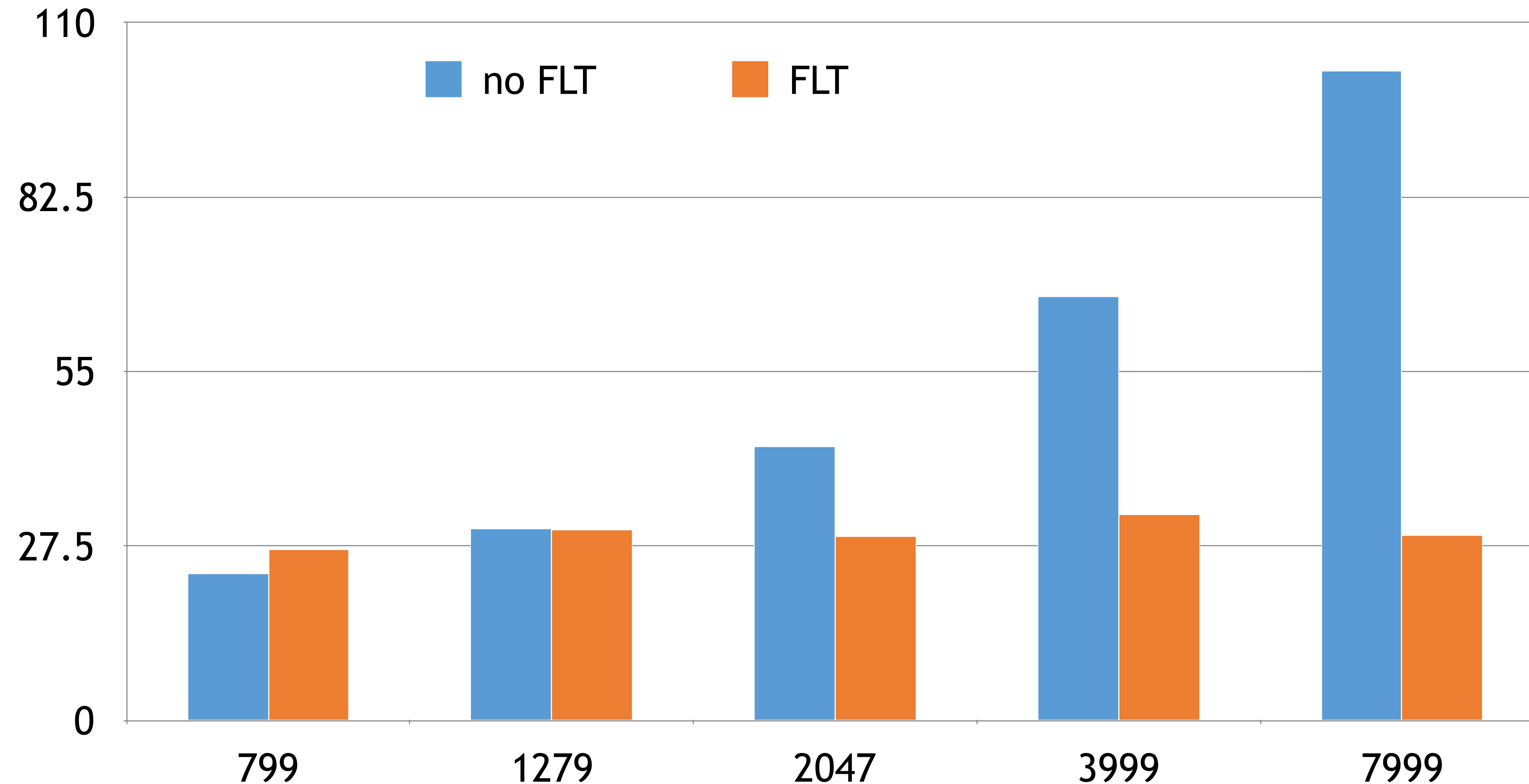




# Fast Legendre Transform

## floating point operations

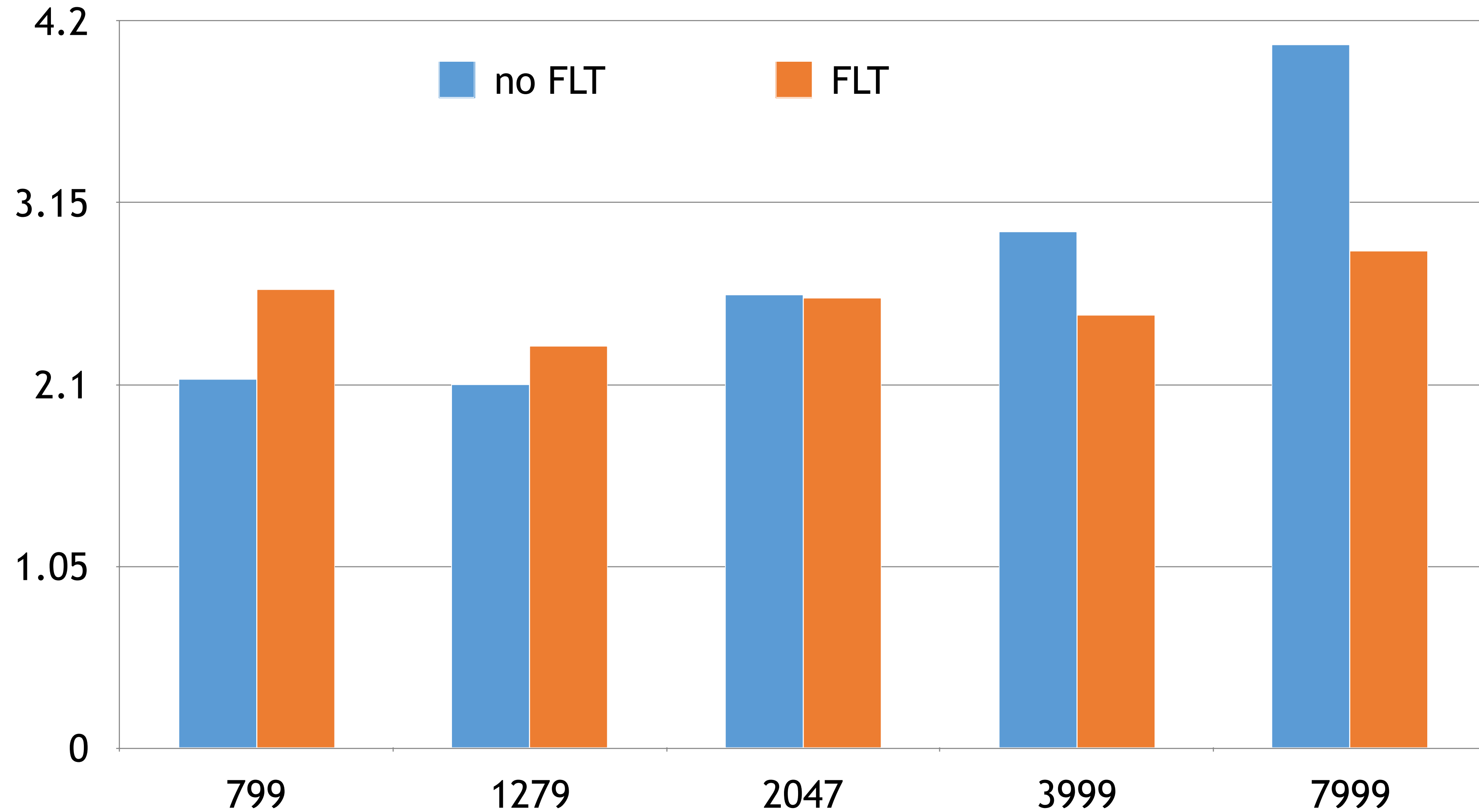
Number of floating point operations for direct or inverse spectral transforms of a single field, scaled by  $N^2 \log^3 N$



# Fast Legendre Transform wallclock time



Average wall-clock time compute cost of  $10^7$  spectral transforms scaled by  $N^2 \log^3 N$





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